

Quo Vadis Nuclear Physics in Europe?

Marek Lewitowicz

- **What is NuPECC?**
- **2017 Long Range Plan in Nuclear Physics**
 - **Nuclear Structure and Physics with Radioactive Ion Beams (RIB)**
 - **RIB facilities**

**The European Expert Board
for Nuclear Physics
associated to European Science
Foundation
*Almost 30 years old***

**Representing
about 6000 scientists**

Members: 31 representatives
from 20 countries + JINR Dubna
5 observers

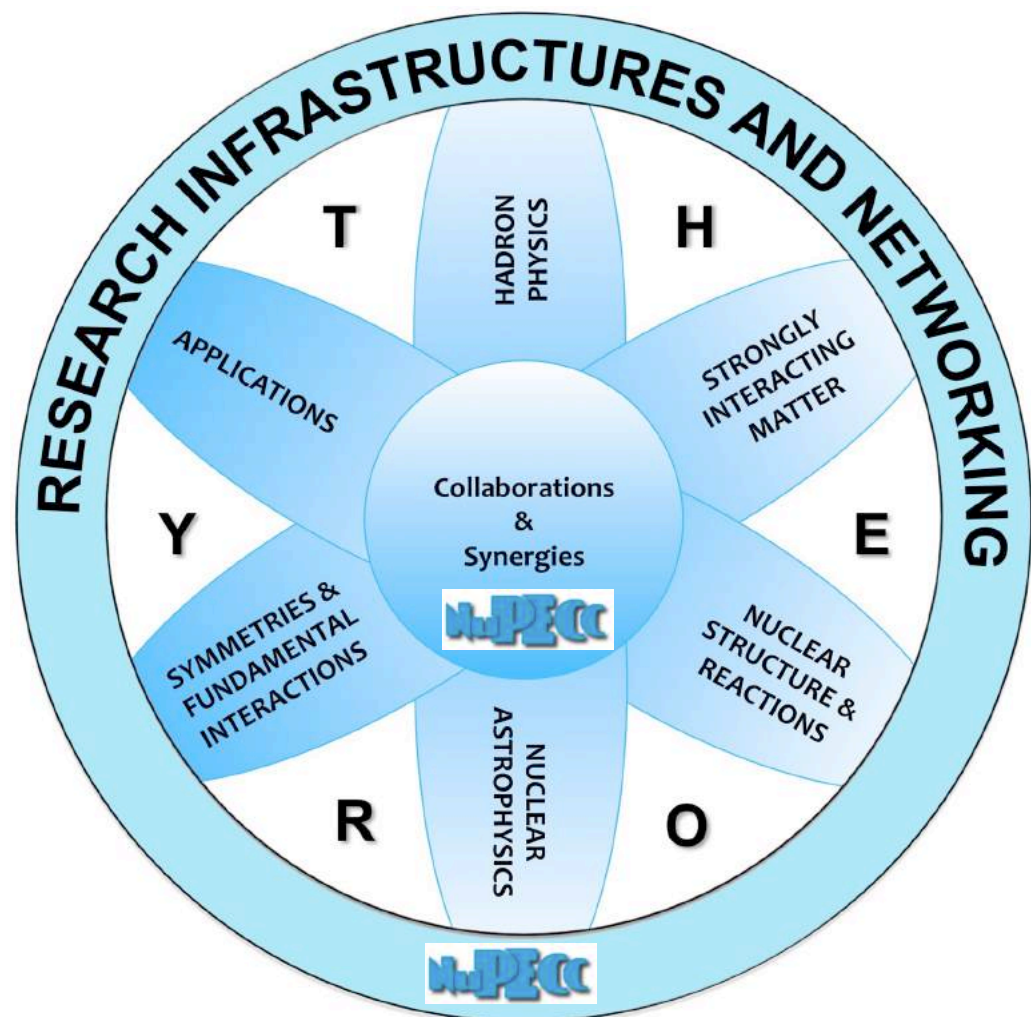
Main mission: elaborate
a strategy at the European scale
for Nuclear Physics



The objective of NuPECC is to:

- develop the strategy for European Collaboration in nuclear science by supporting collaborative ventures between research groups within Europe -> Long Range Plan
- promote nuclear physics and its trans-disciplinary use in applications for societal benefit.

NuPECC ToR

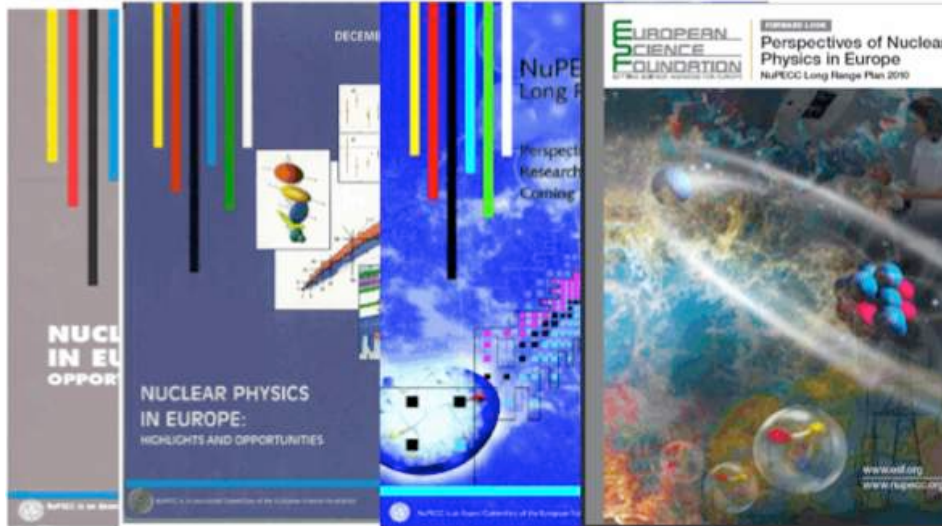


1991

1997

2004

2010



2017

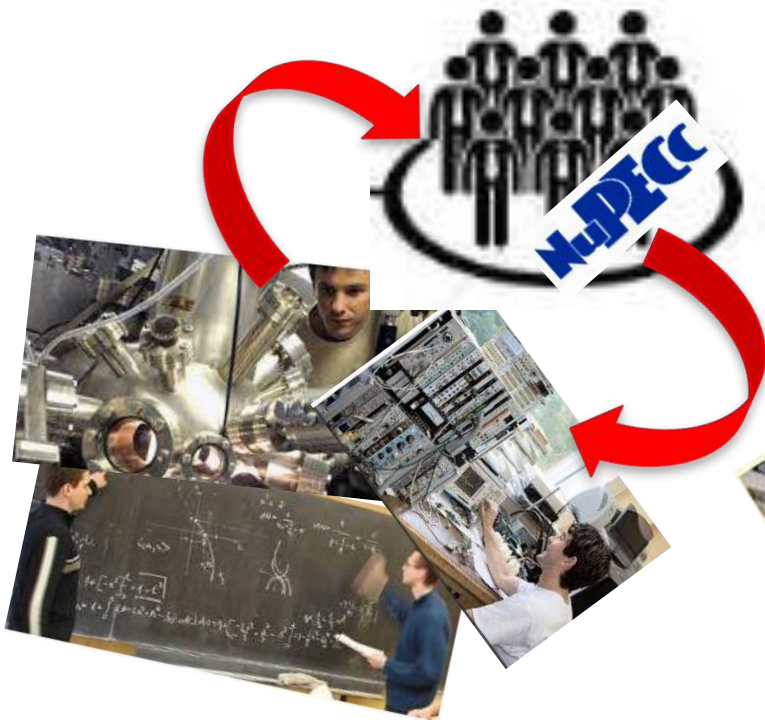


- The LRP identifies opportunities and priorities for the nuclear science in Europe
- The LRP provides the European Commission and national funding agencies with a framework for coordinated advances in nuclear science in Europe

<http://www.nupecc.org/lrp2016/Documents/lrp2017.pdf>

Who did produce this strategic document ?

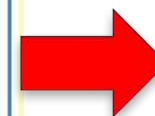
<http://www.nupecc.org/lrp2016/Documents/lrp2017.pdf>



**Community working in the field :
200 experts in experiments and theory
7 working groups**



**Town meeting
in Darmstadt
January 2017**



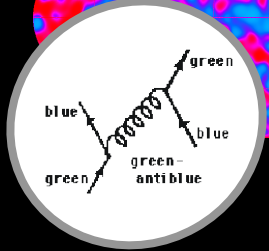
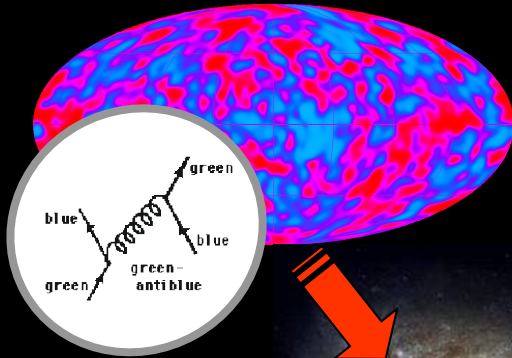
**Report
June 2017**

**2016
beginning**

end 2017

LRP presentation – Brussels November 27, 2017

The origin of life



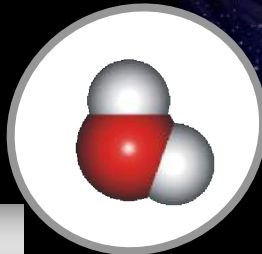
Creation of Matter

High Energy Physics



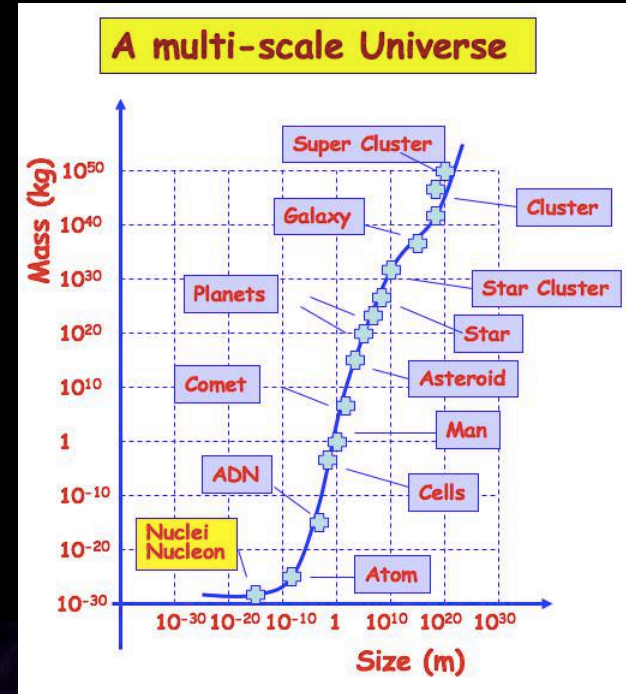
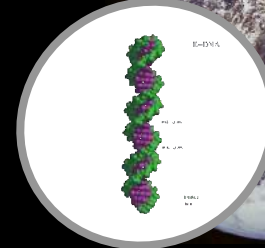
Nuclear Physics

Elements



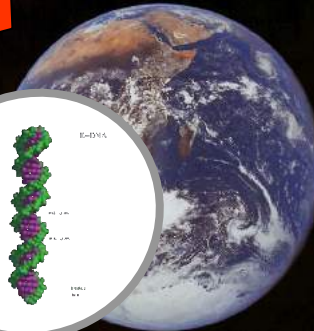
Chemistry

Biochemistry



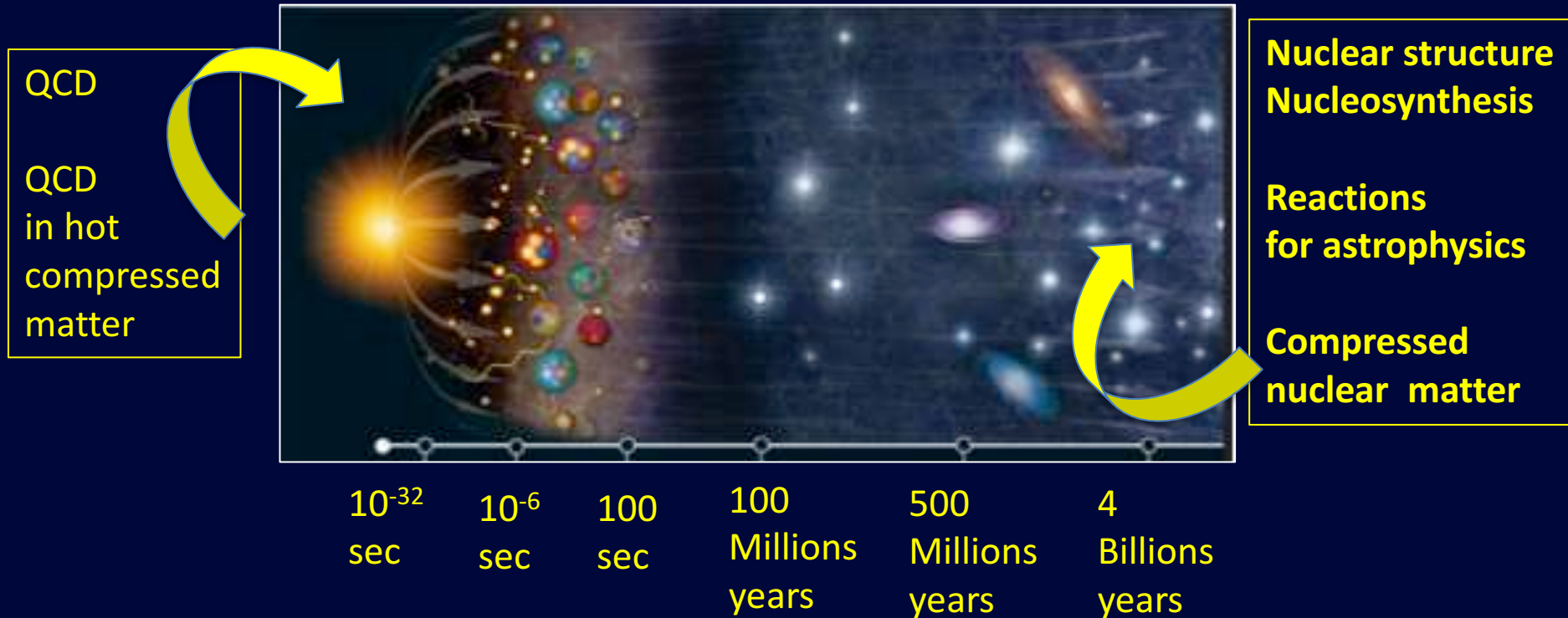
Planets

Life



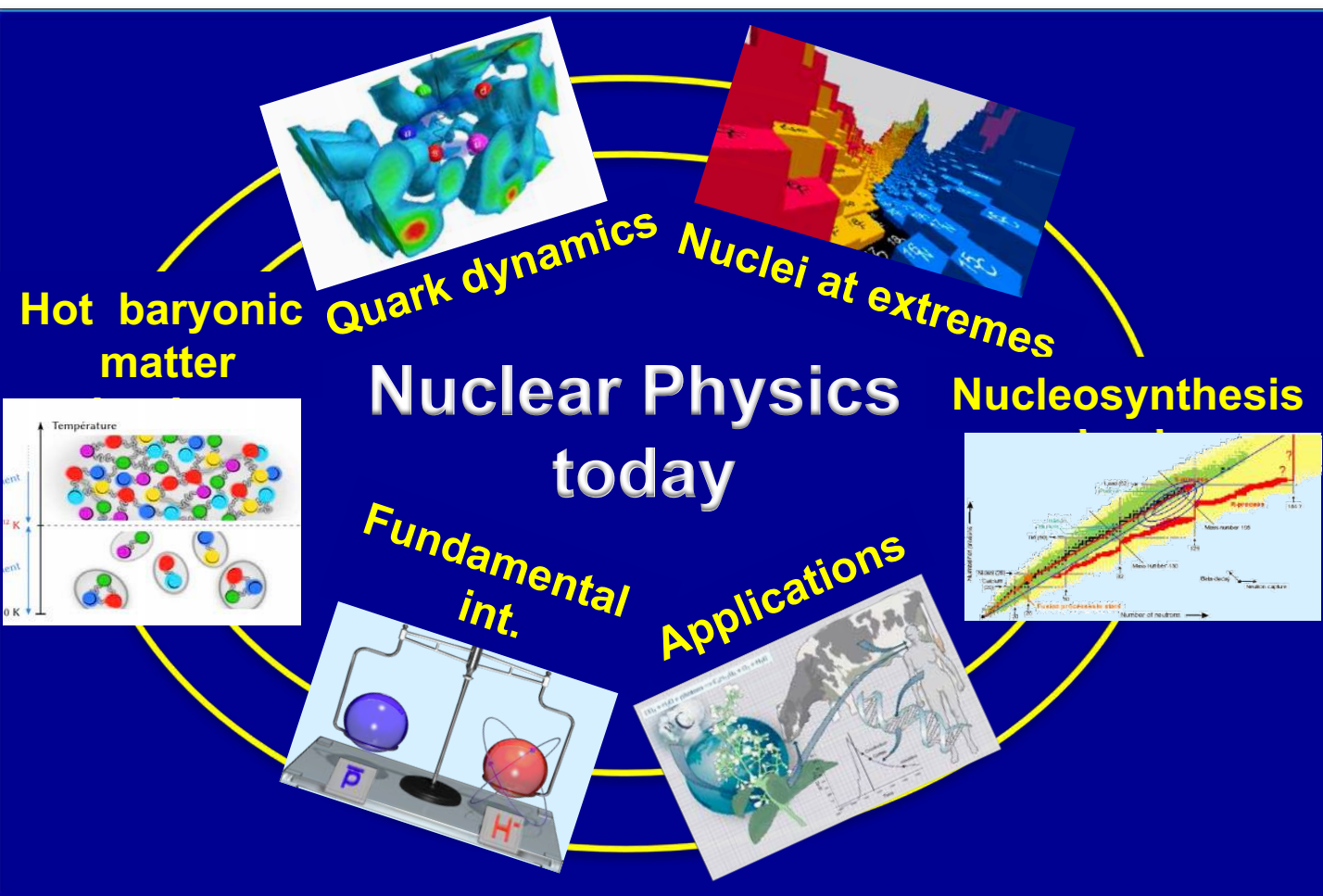
Nuclear physics and the evolution of the Universe

Nuclear Physics with its different research domains addresses several key issues for the understanding of the different stages of the evolution of the universe



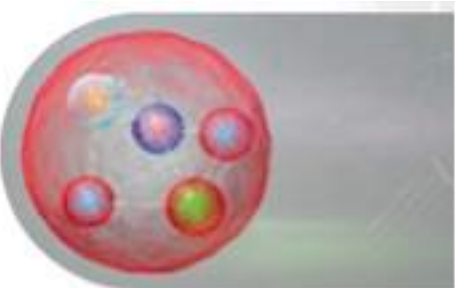
To tackle the different problems one needs a distributed approach and efforts :
different accelerator types and energies

**Study of nuclear matter in all its forms
exploring their possible applications**



**Nuclear physics
is a very broad
field !**

**Each area
needs
particular
tools and
technologies**



Hadron physics involves studying the intricate patterns of the hadron structures that emerge, the spectra mapping their energies and their interaction



- How is mass generated in Quantum Chromo Dynamics (QCD) and what are the static and dynamical properties of hadrons?

- How does the strong force emerge from the underlying quark-gluon structure of nucleons?

Identify the connection between quark dynamics and quantum numbers (spin and orbital angular momentum)



Experiments



Theory

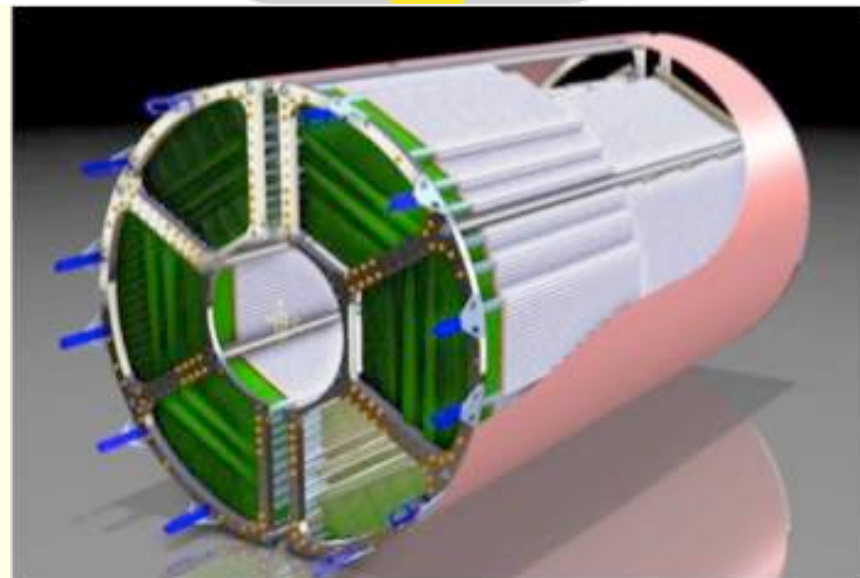
QCD :
theory of
strong
force



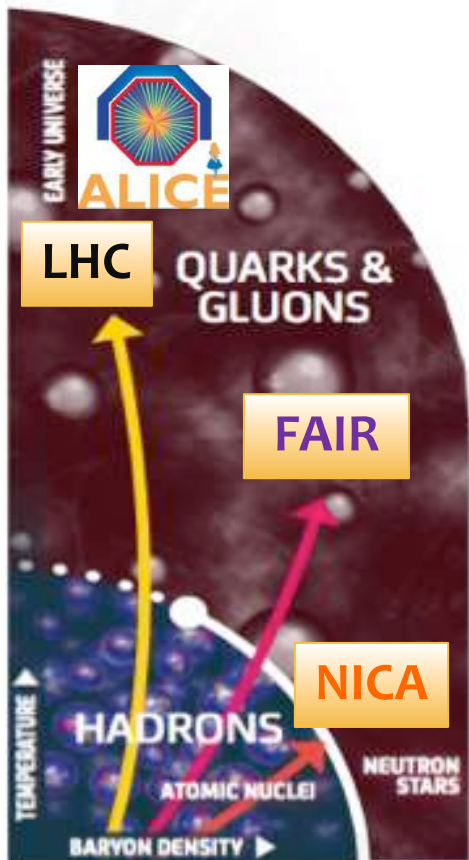
The proton

Studies have uncovered discrepancies in measurements of the proton radius made with different techniques.

New experiment planned to tackle this issue (one at Mainz-MESA) : new physics?



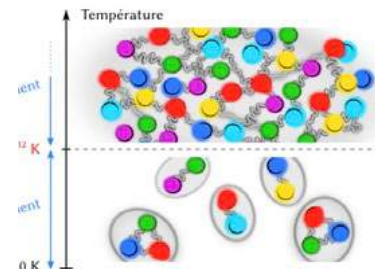
High resolution experiments with antiprotons (PANDA) at FAIR will address many issues to test in detail theory of Quantum Chromo Dynamics (QCD)



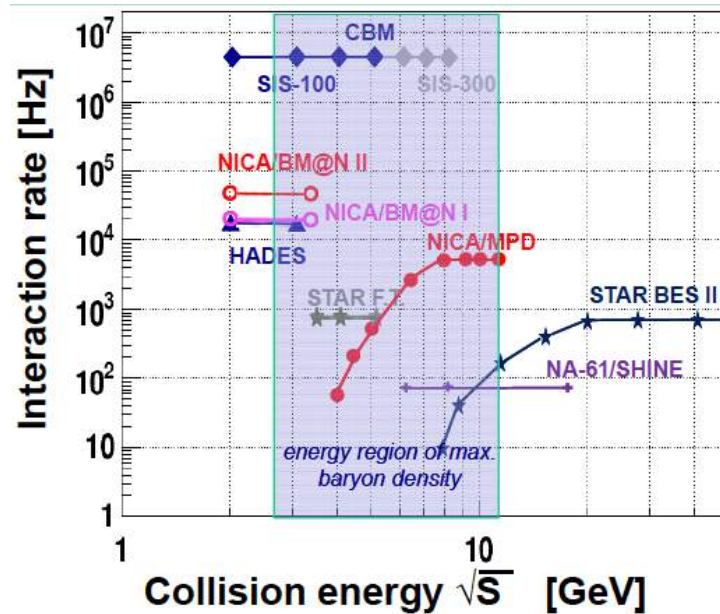
Matter at very high temperature and density Quark Gluon Plasma (QGP) reveals the high energy processes that drove the evolution of the universe after its birth.

Its very exotic nature is found in massively compressed stellar corpses : neutron stars

QGP turned into hadrons few μs after Big Bang. It is not seen in astronomical observations and thus is recreated in the Heavy Ion labs within volumes of nuclear size



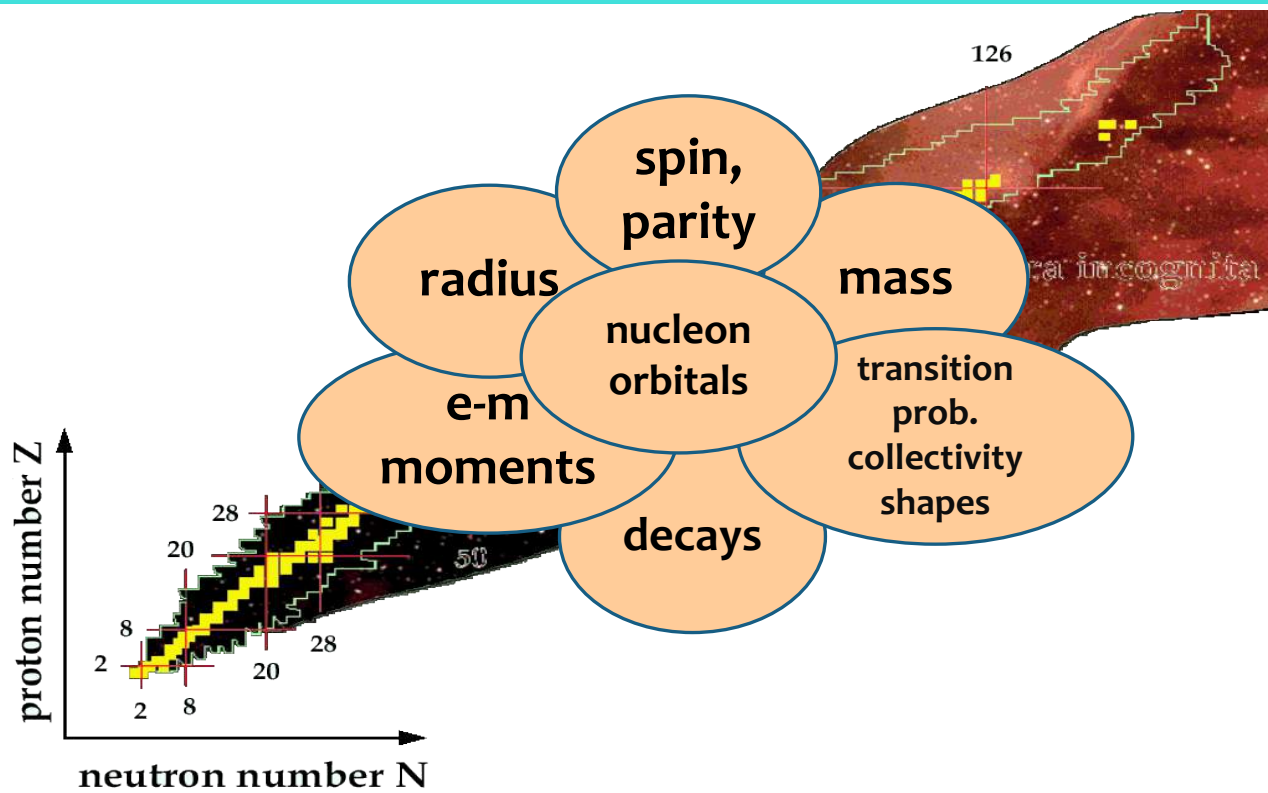
signals from de-confinement and chiral symmetry restoration



Questions

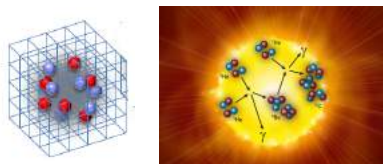
- Where are the limits of stability and what is the heaviest element?
- How does nuclear structure evolve (also with T and L) and what shapes can nuclei adopt ?
- How complex are nuclear excitations?
- How do correlations appear in dilute neutron matter ?
- What is the density and isospin dependence of the nuclear equation of state ? (for neutron stars)

Protons and neutrons via the nuclear force can create as many as 7000 nuclear species (250 stable, about 2000 synthesised). At limit of stability there is an extreme proportion of p or n in a nucleus.



Tools : Heavy ion beams far from stability produced with different reaction types, gamma beams and radioactif decay

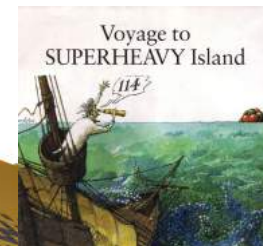
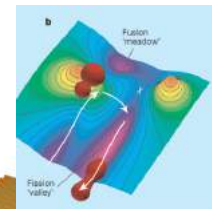
Lattice Effective Field Theory



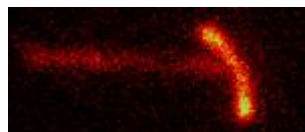
Equation of state



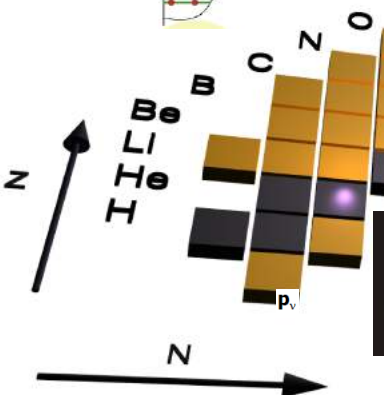
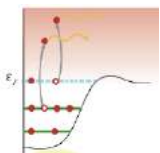
Fission dynamics



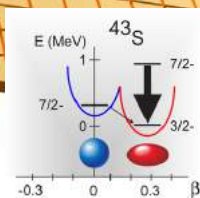
1p, 2p radioactivity



Coupling to continuum

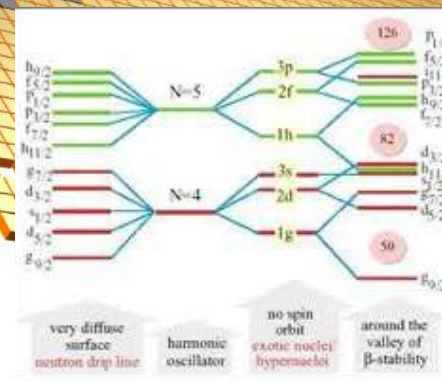


Clusters

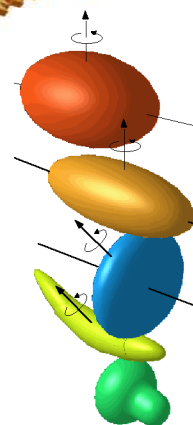


Shape Coexistence

Limits of existence



New magic numbers



Exotic Shapes



Neutron halos

Search and UNDERSTAND regular and simple patterns in the structure of complex nuclei by characterizing them under EXTREME conditions (E^*, J, T):

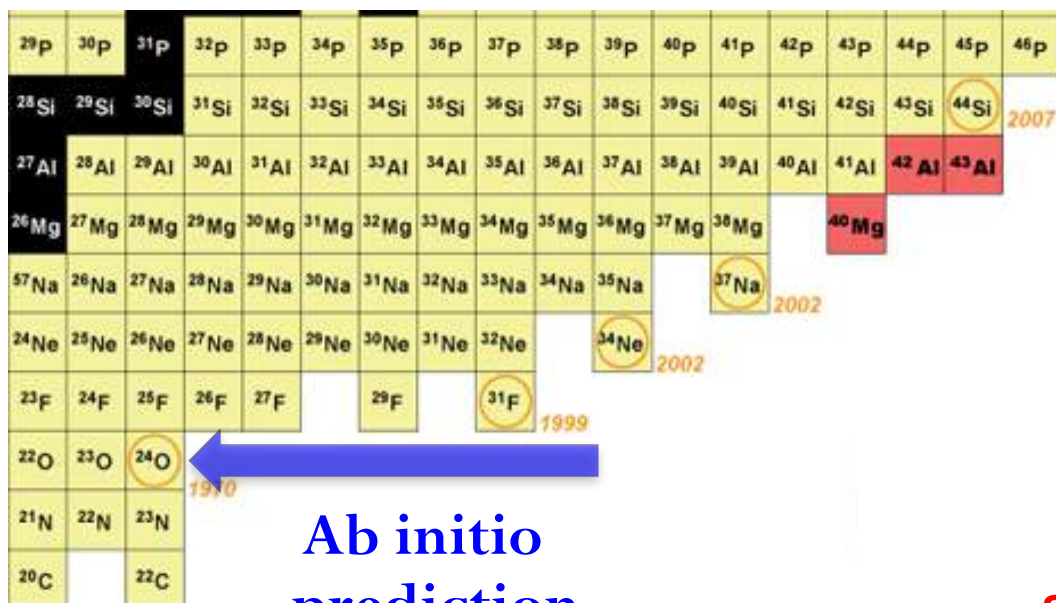
amplify different aspects of the interaction

Where (and what) is the nuclear dripline?

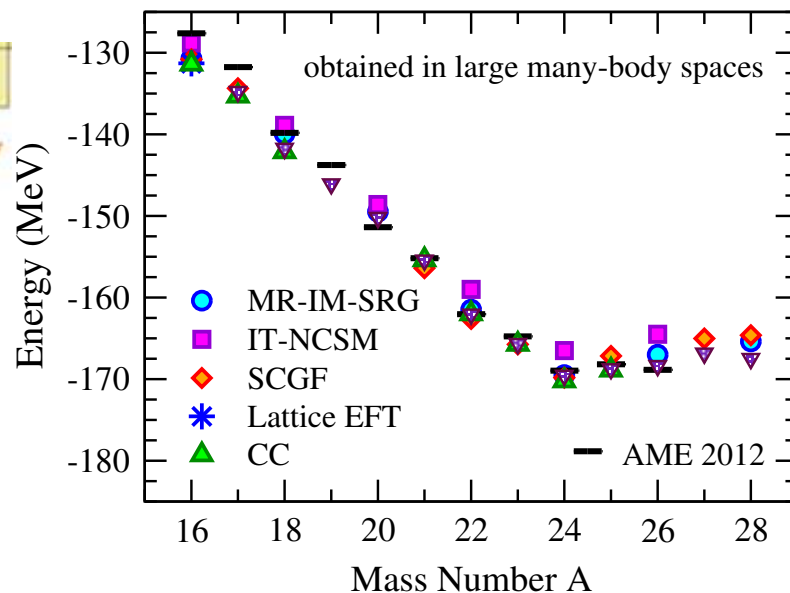
Limits defined as last isotope with positive neutron separation energy

- Nucleons “drip” out of nucleus

Neutron dripline experimentally established to Z=8



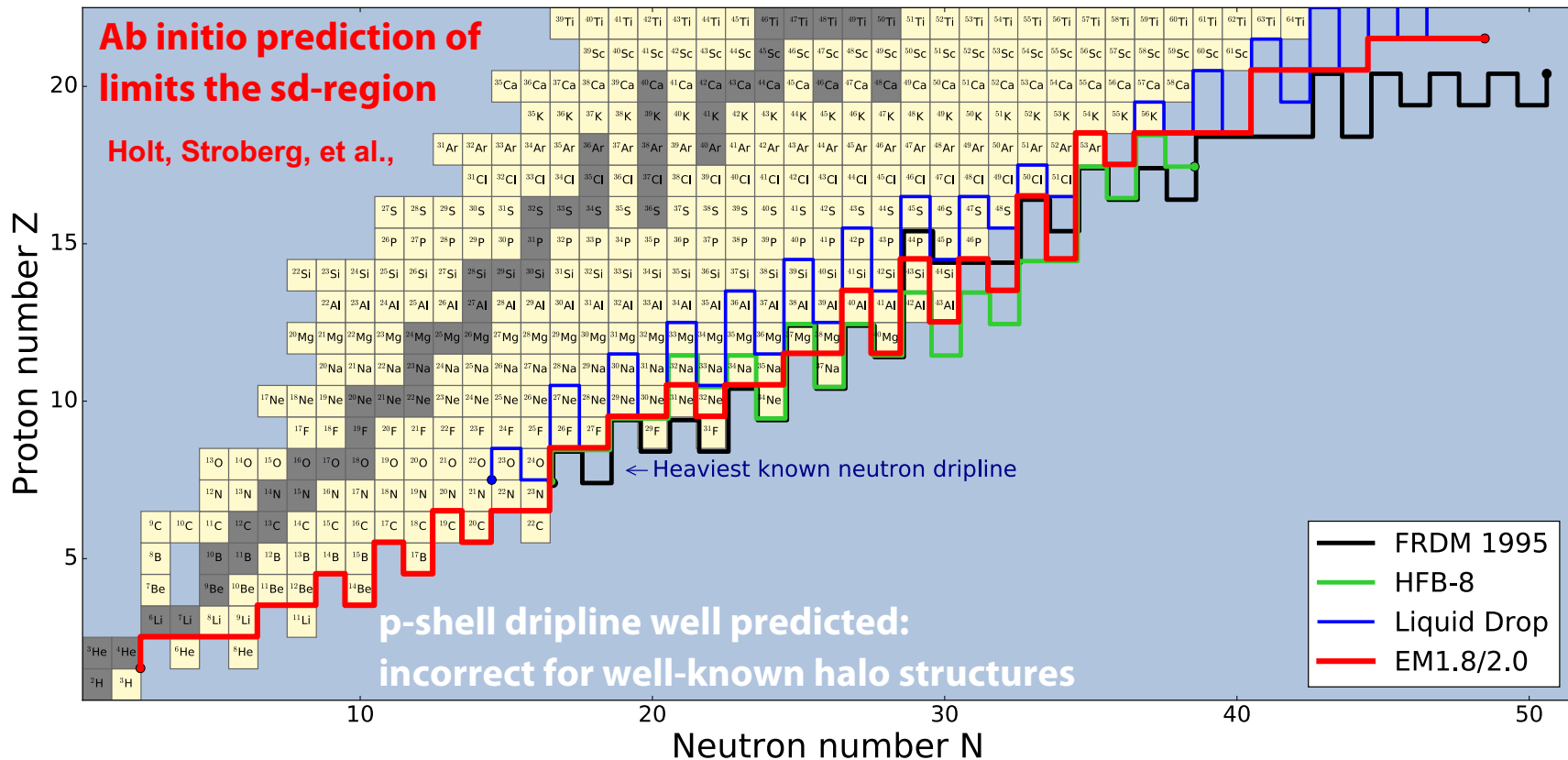
Ab initio prediction (2016)



Same result from same input NN+3N forces

Already well beyond where fit to data!

Courtesy of J. Holt



General agreement with FRDM/HFB predictions (not available for p-shell)

Very strong experimental programs at and competition from RIKEN, FRIB, FAIR, RAON !

Courtesy of J. Holt

Shell evolution novel phenomenon in exotic nuclei

Unexpected magic numbers appear

Expected ones disappear

Can we understand/predict this trend?

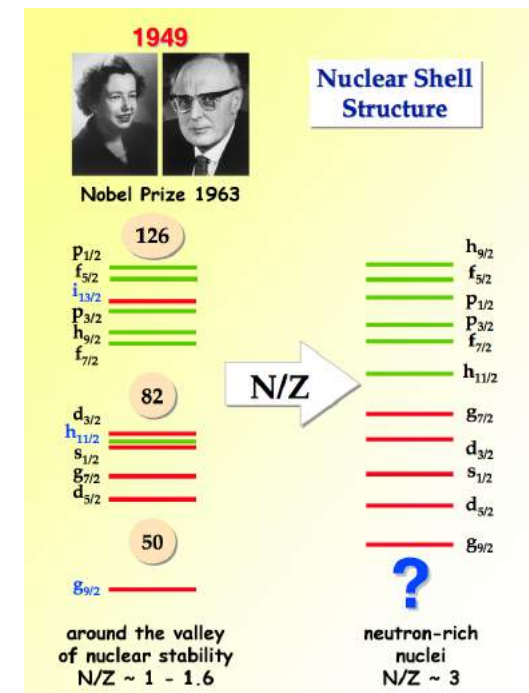
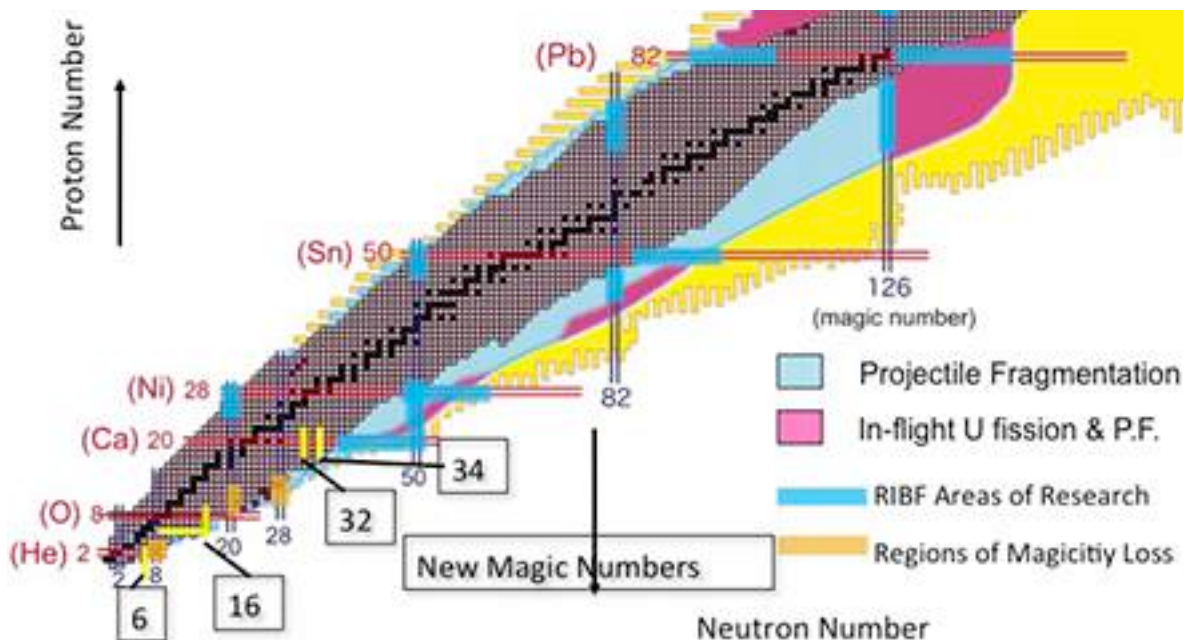
Signatures of Magic Numbers

Sudden decrease in separation energies (masses)

Elevated first excited state (spectroscopy)

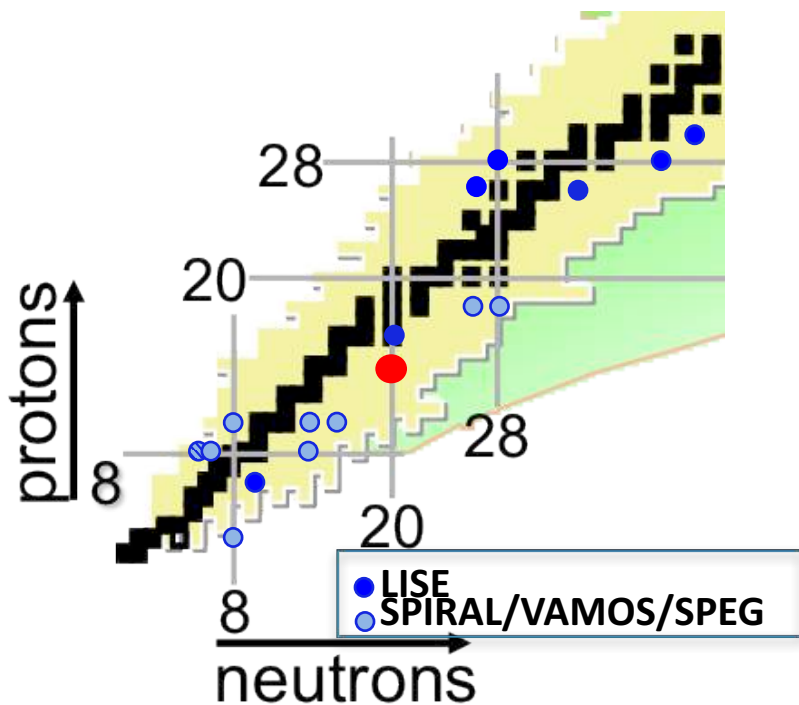
Tightly bound (decreased radii)

Must observe all signatures – many experiments needed!

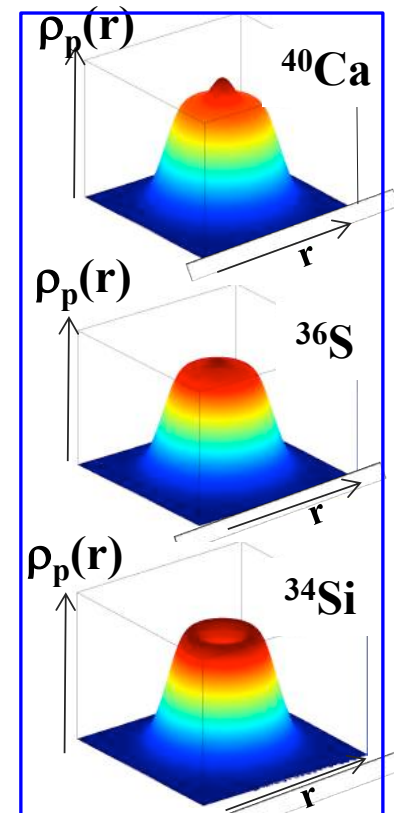
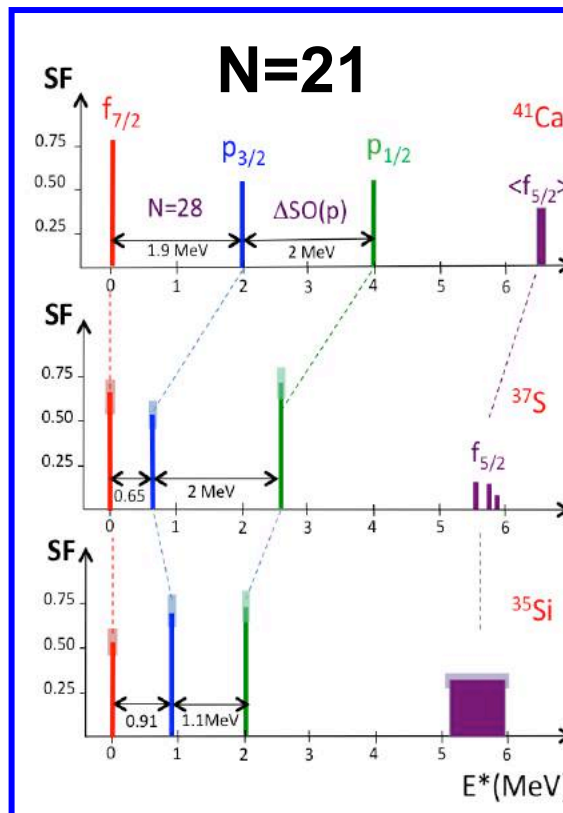


Courtesy of J. Holt

- Probe single particle energies, shape coexistence or pairing modes via $^{34}\text{Si}(d,p)$, (d,t) , $(d,^3\text{He})$, (p,d) , (p,t) , $(p,^3\text{He})$... reactions.
- Variety of beams at **10-20MeV/A**
- Instrumentation: magnetic spectrometers, Si and gamma arrays



Spin - orbit interaction in the ^{34}Si bubble nucleus



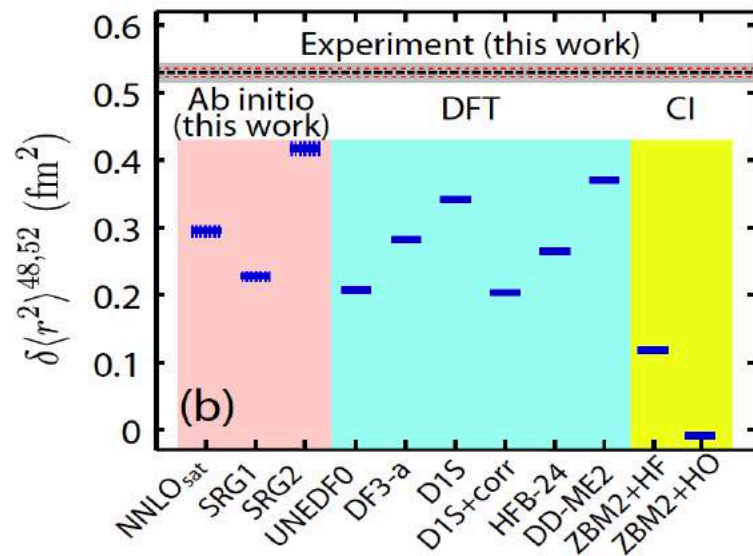
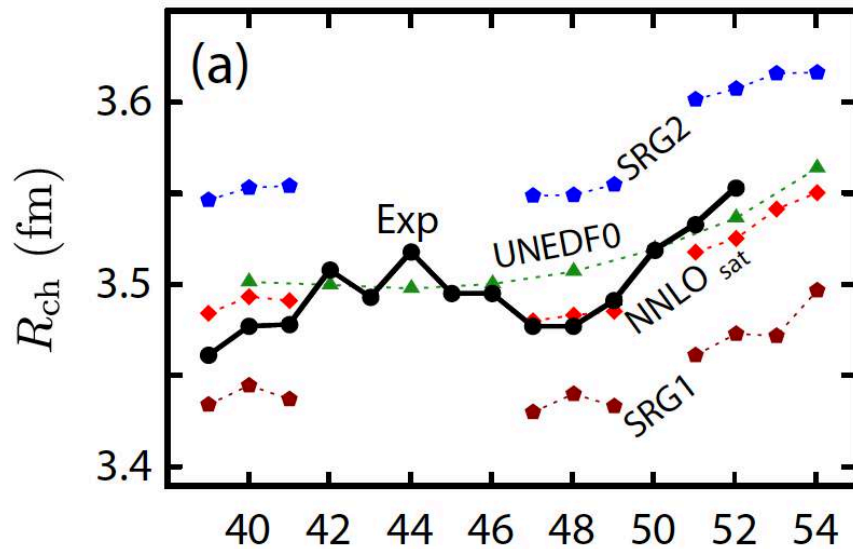
G. Burgunder et al. PRL (2014)

Charge radii of $^{49,51,52}\text{Ca}$, obtained from laser spectroscopy experiments at ISOLDE, CERN

Unexpected large charge radius questions the magicity of ^{52}Ca

Theoretical models all underestimate the charge radius

Ab-initio calculations reproduce the trend of charge radii



R. F. Garcia Ruiz *et al*, Nature Physics (2016)

High-sensitivity for nuclear structure of exotic nuclei – used in several EU laboratories



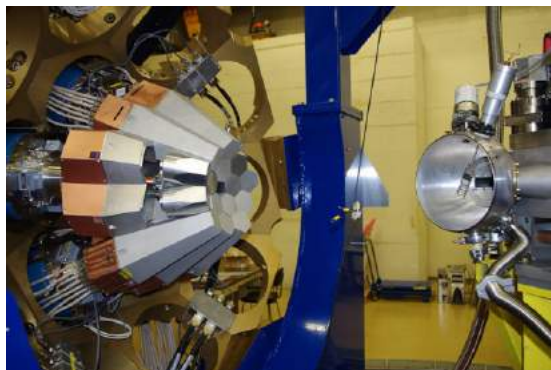
2010 → 2011 LNL, Italy
5TC (15 detectors)



2012 → GSI, Germany
6TC+3 DC (22 detectors)



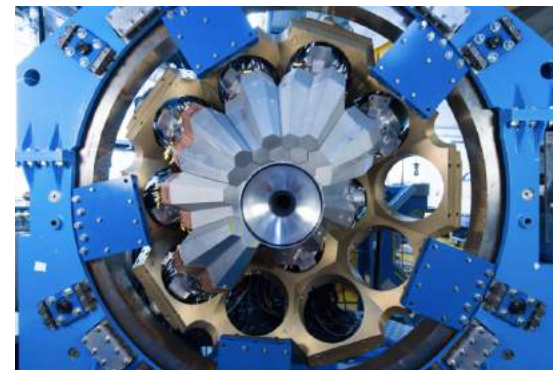
2014 → GANIL, France
15TC (45 detectors)



AGATA Demo.+PRISMA
Total Eff_{Nominal} ~2.6%



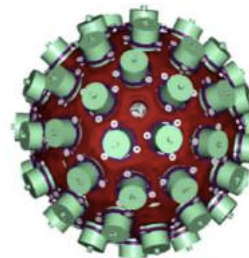
AGATA @ FRS
Total Eff. ($\beta=0.5$) ~ 10%



AGATA @G1
Total Eff ~ 8% to 14%

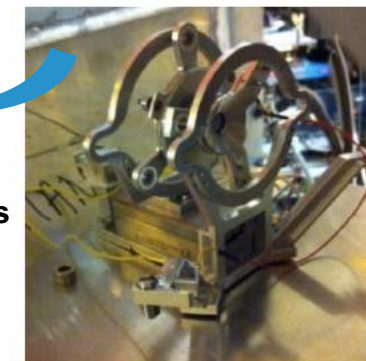
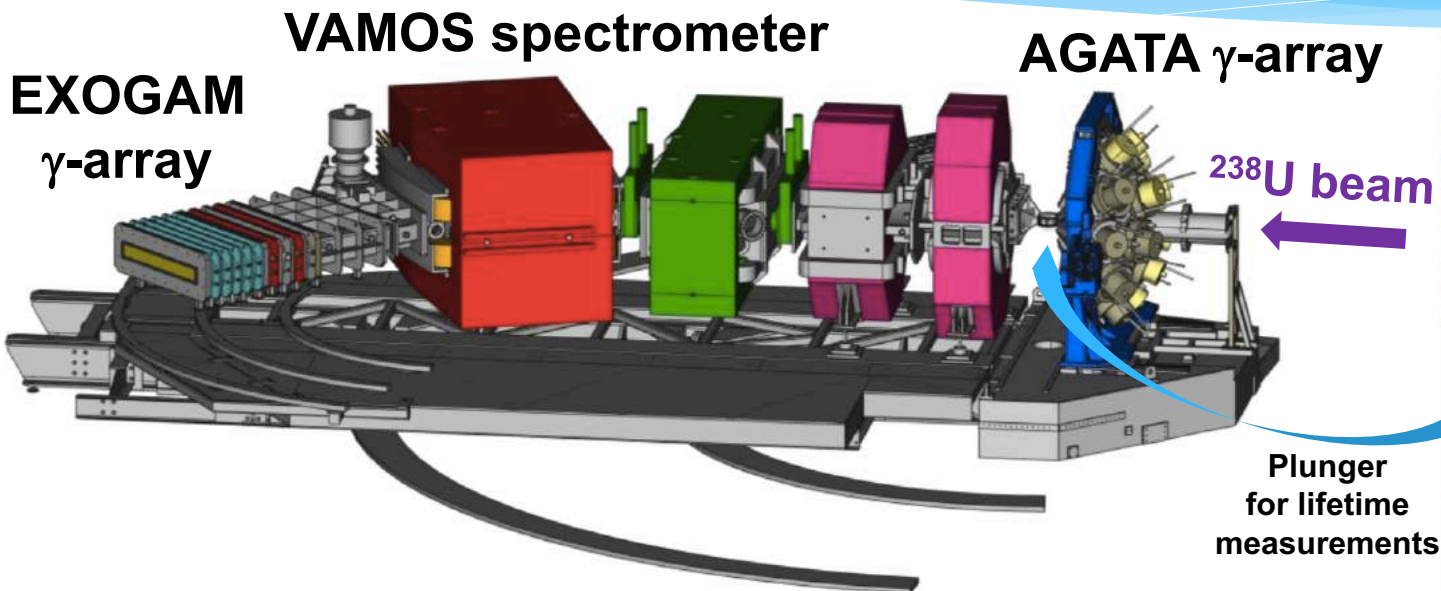
AGATA array: A powerful traveling instrument - its construction has to proceed in the next years up to 4π coverage (60 triple clusters = 160 detectors) !

AGATA 4π

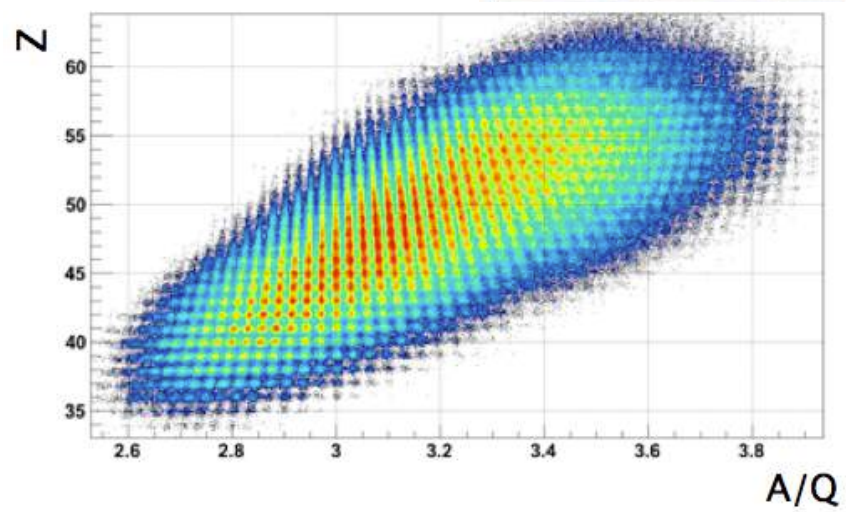
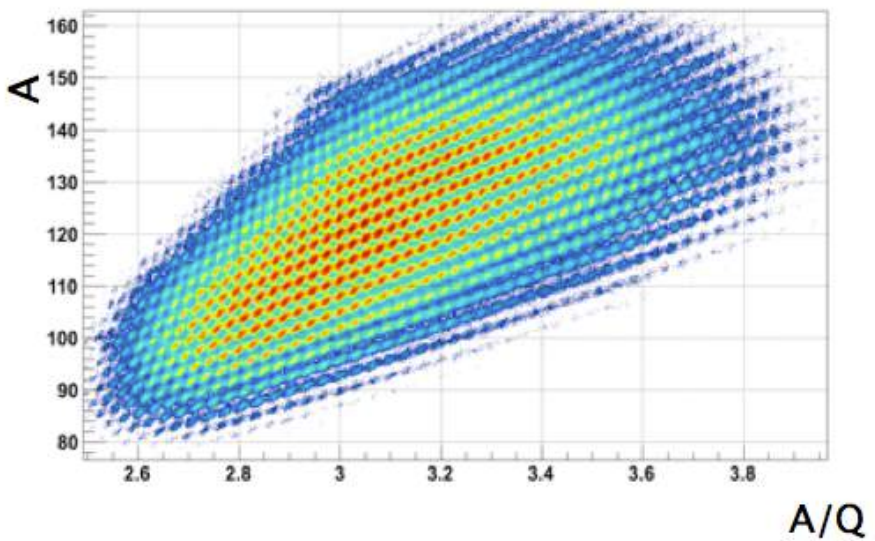


Tripple Cluster





Full identification of produced isotopes in Z, Q & A



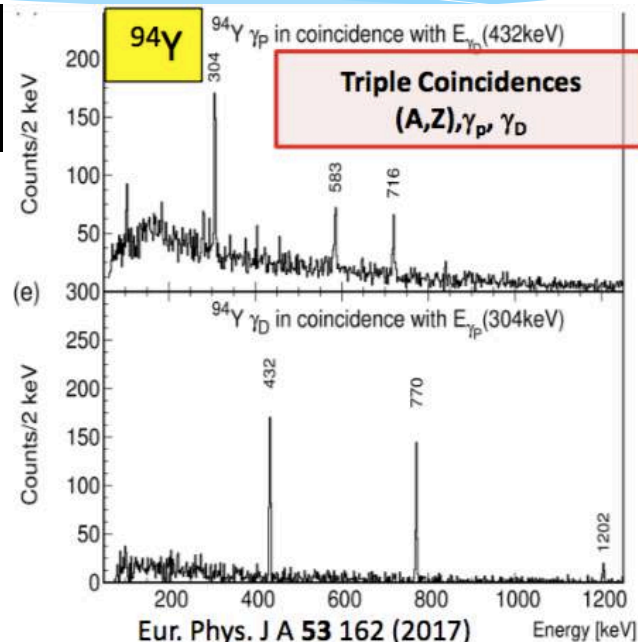
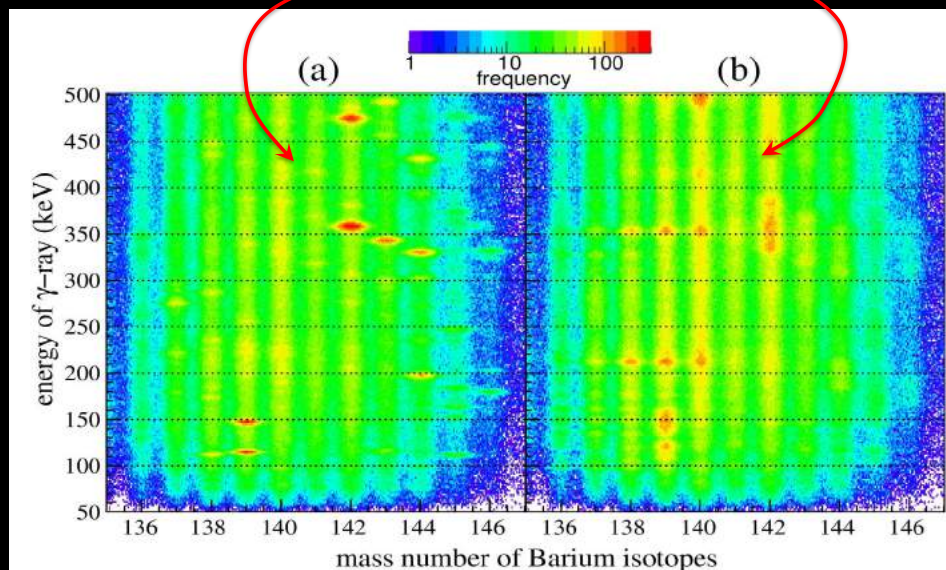


Ba (Z=56)
in VAMOS

Depending on Doppler correction
 γ -rays of Barium or Zirconium



Zr (Z=40)



Eur. Phys. J A 53 162 (2017)

$2^+ \rightarrow 0^+$

$4^+ \rightarrow 2^+$

120 μm

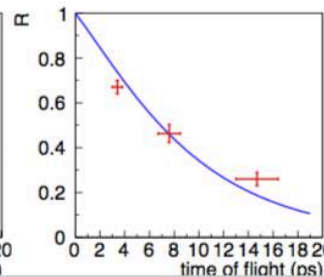
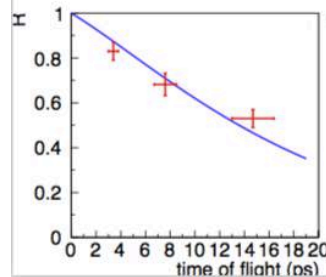
270 μm

520 μm

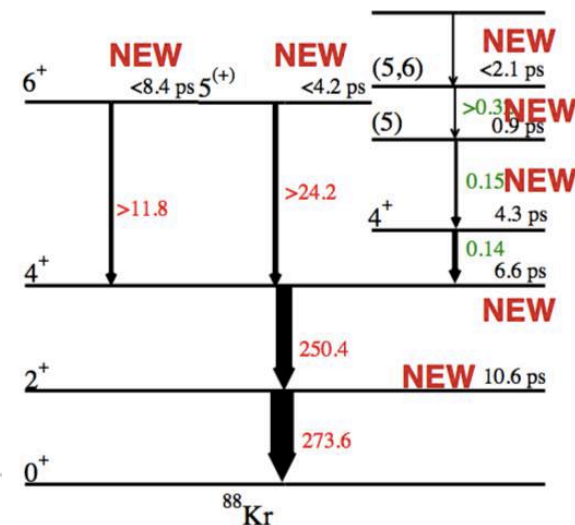
Lifetime measurements: ^{88}Kr

$2^+ \rightarrow 0^+$

$4^+ \rightarrow 2^+$



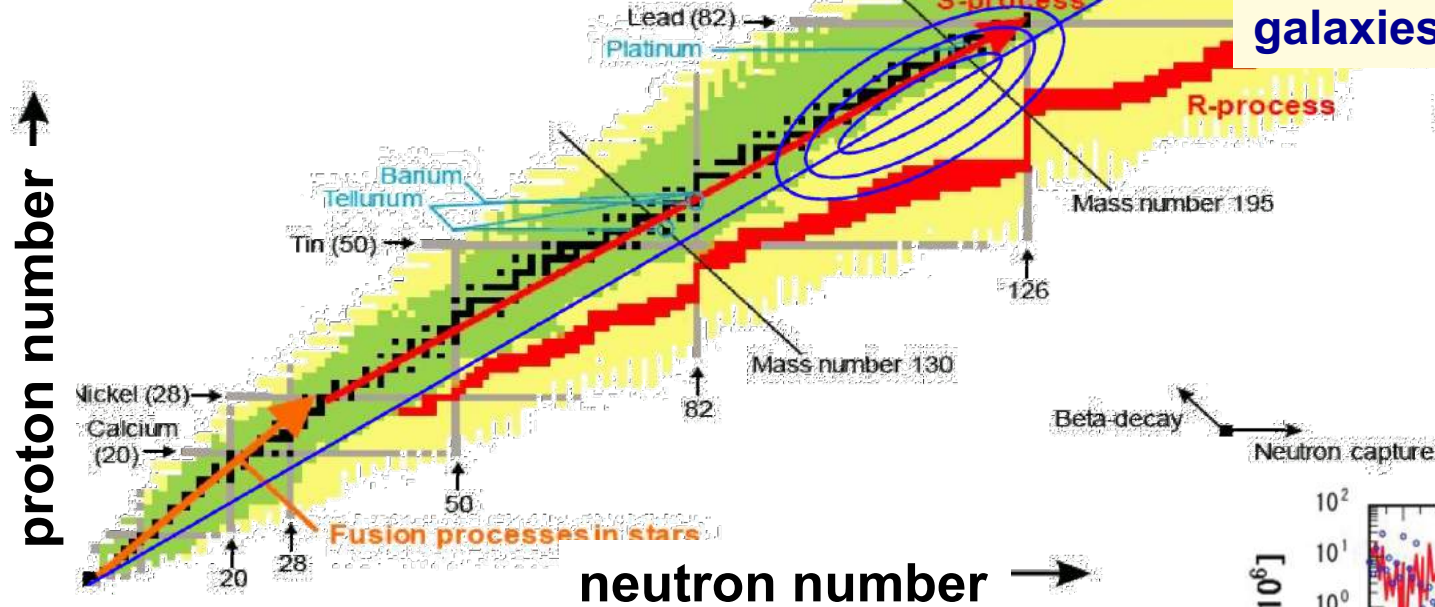
$B(E2) \downarrow (e^2\text{fm}^4)$
 $B(M1) \downarrow (\mu_N^2)$



Various nucleosynthesis processes

- BBN
- Fusion processes in stars
- Explosive nucleosynthesis

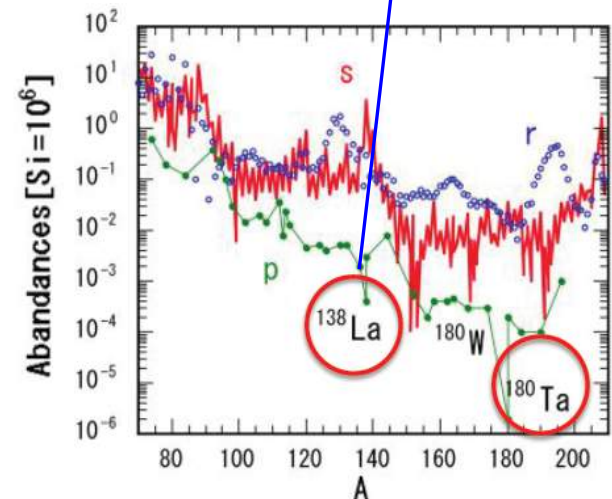
What are the nuclear processes that drive the evolution of the stars, galaxies and the Universe?



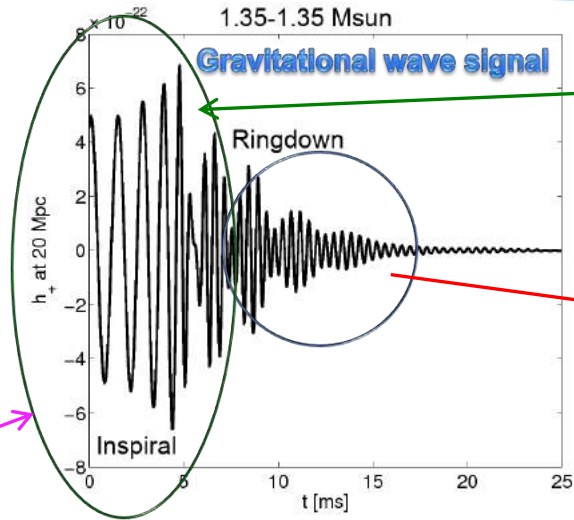
Interplay of:

- nuclear structure
- nuclear decays
- half-lives
- nuclear reactions
- nuclear masses

Proton and neutron rich side nuclei – our precursors need to be investigated with different tools



Neutron star mergers: Gravitational waves and production of heavy elements



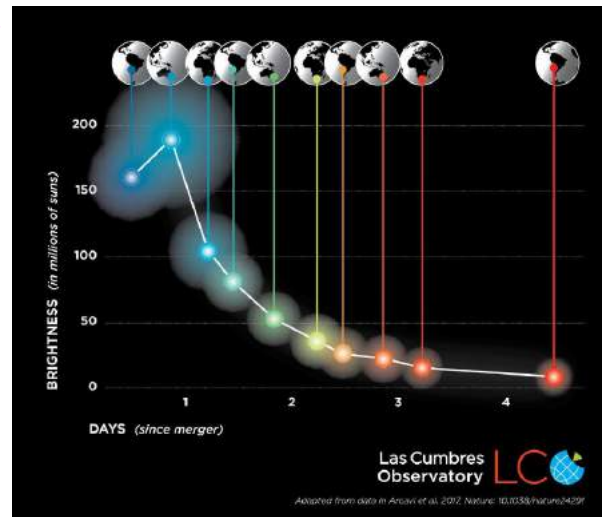
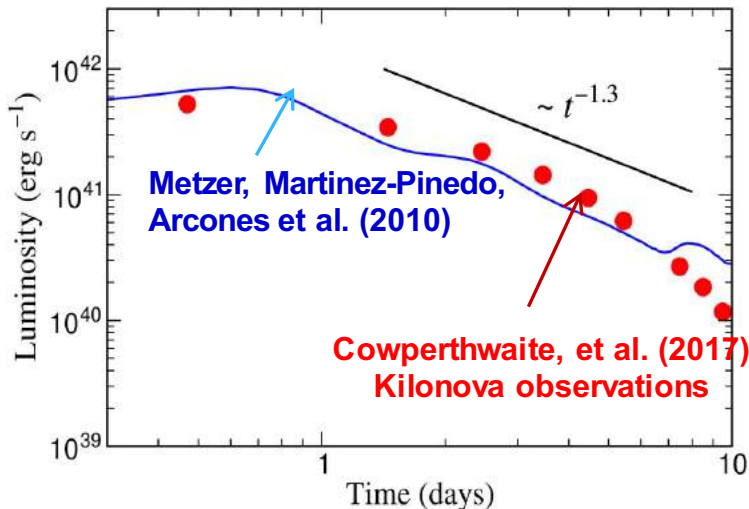
Neutron star mass

This depends on the Nuclear Equation of state

The messengers from neutron star mergers :

- **Gravitational waves**
- **Electromagnetic signals** characterizing the nuclei in the ejecta
- **neutrinos**

Gravitational wave emission seen together with electromagnetic signals

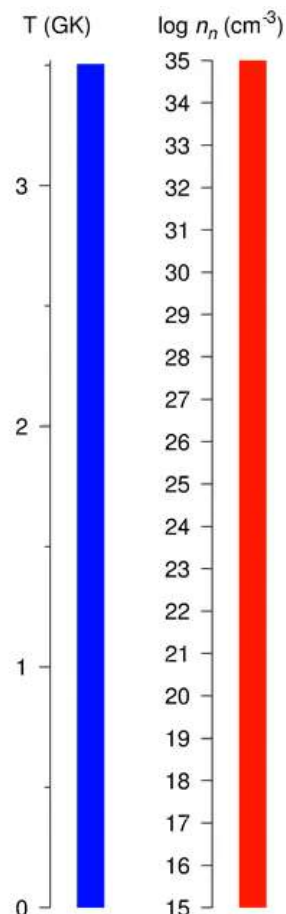
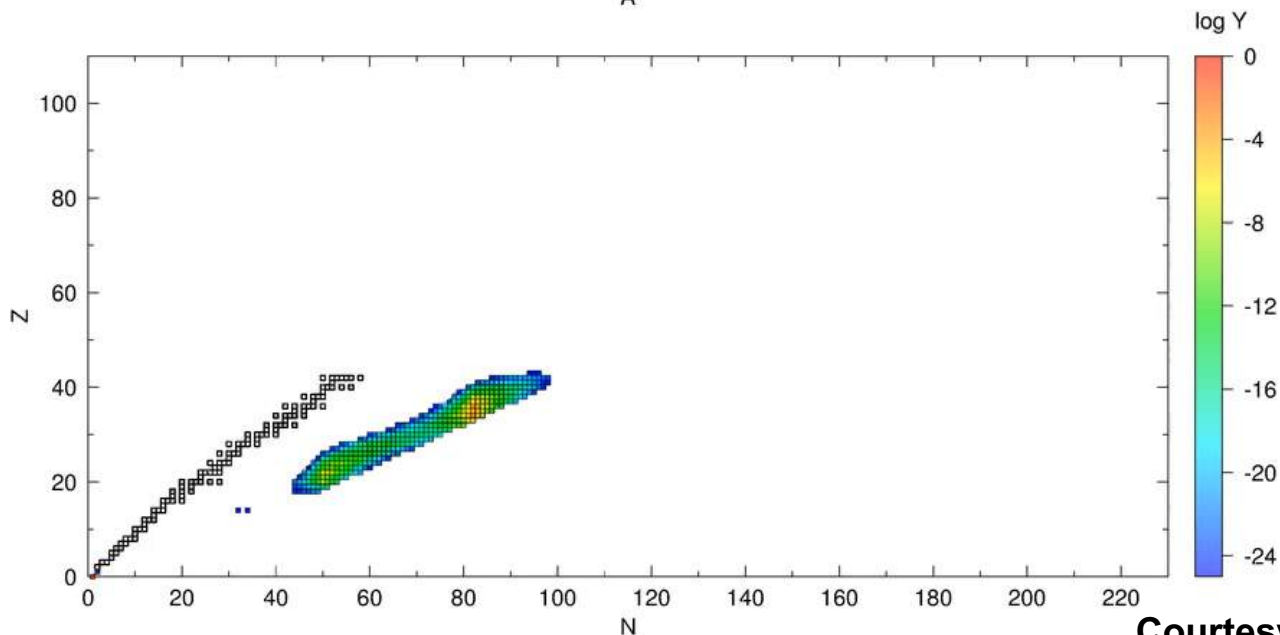
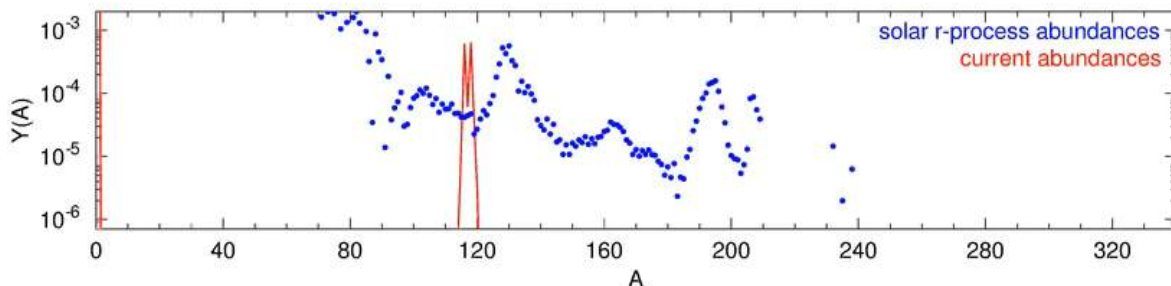


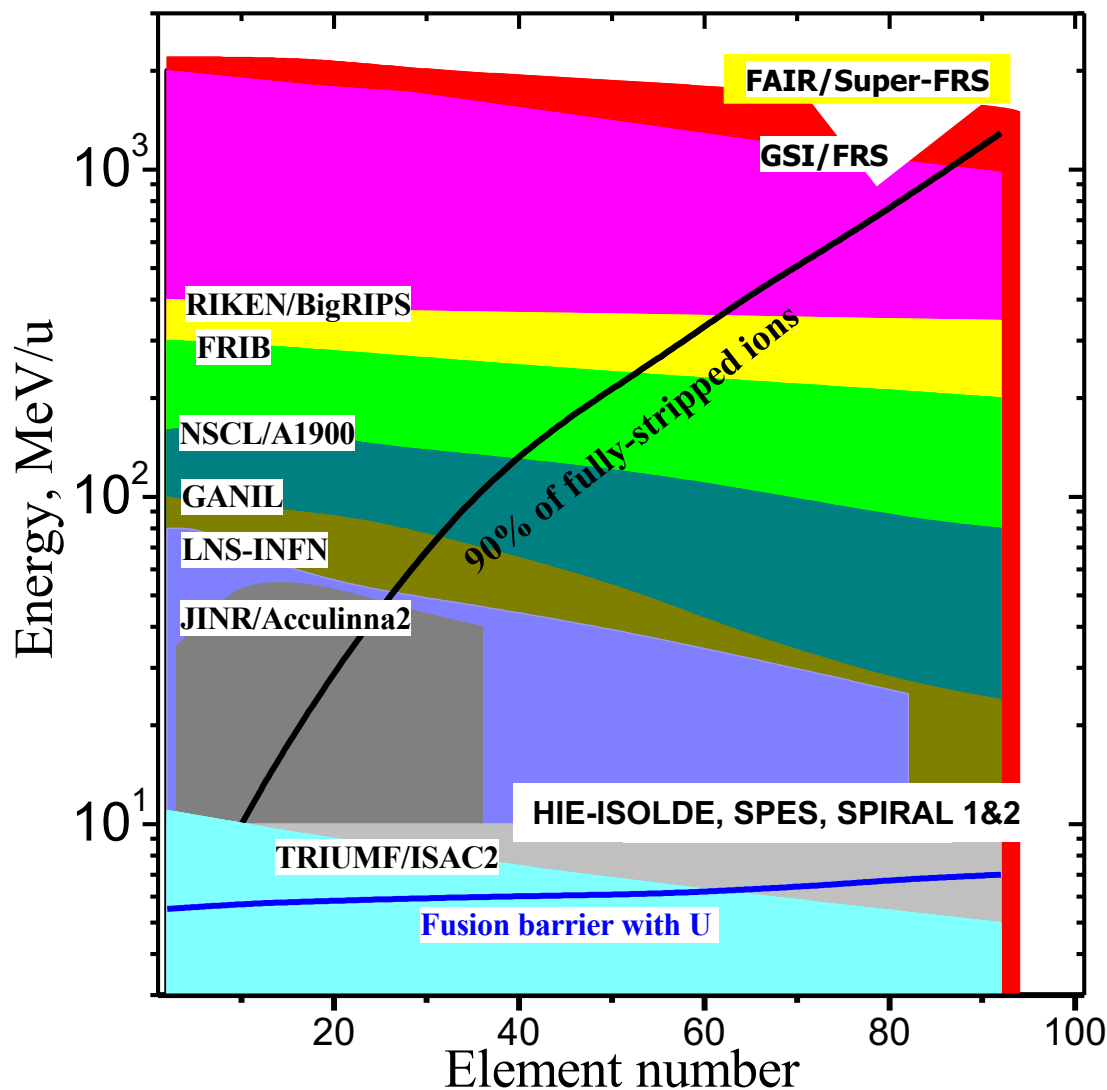
Time evolution determined by the radioactive decay of r-process nuclei (science drive of facilities with RIB)



... a simulation of the r-process

$T = 3.50$ GK, $n_n = 2.937e+35$ cm⁻³, $R_{n/s} = 623.3$, $s = 0.621$ k_B/nuc, $t = 0.0131$ s

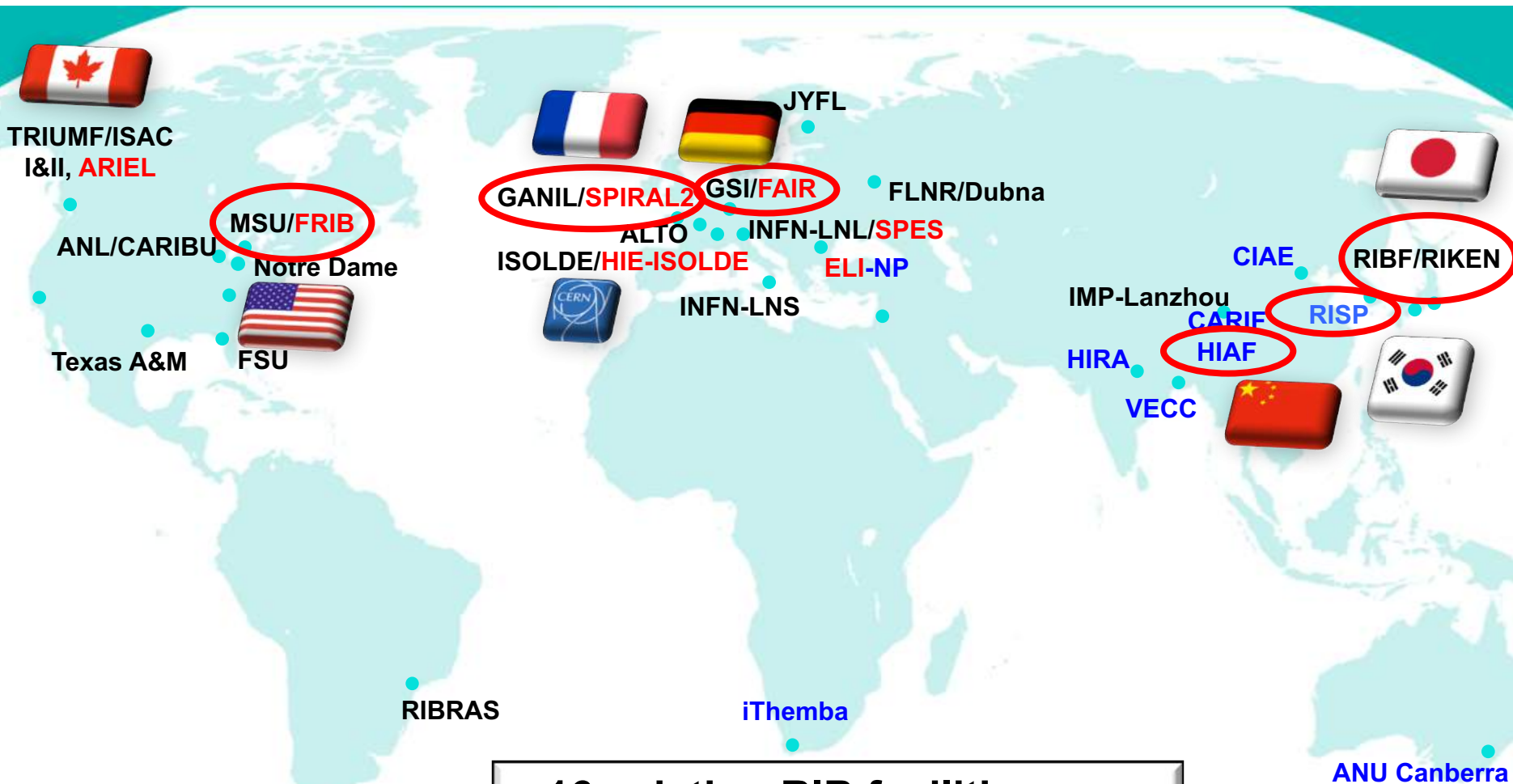




Energy range
from 0 to 2 GeV/n

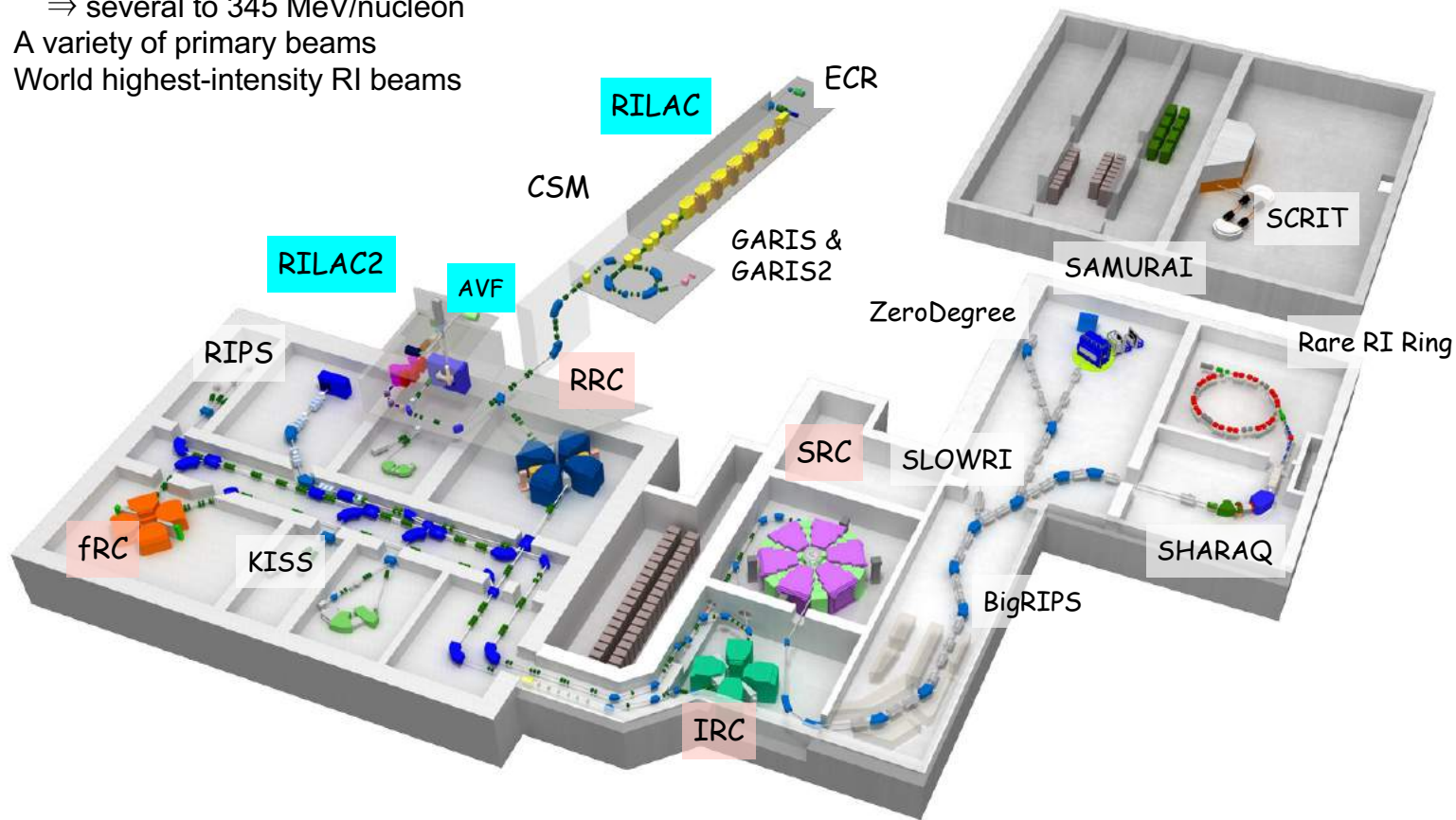
ISOL facilities:
from 0 to 20 MeV/n

Facilities worldwide

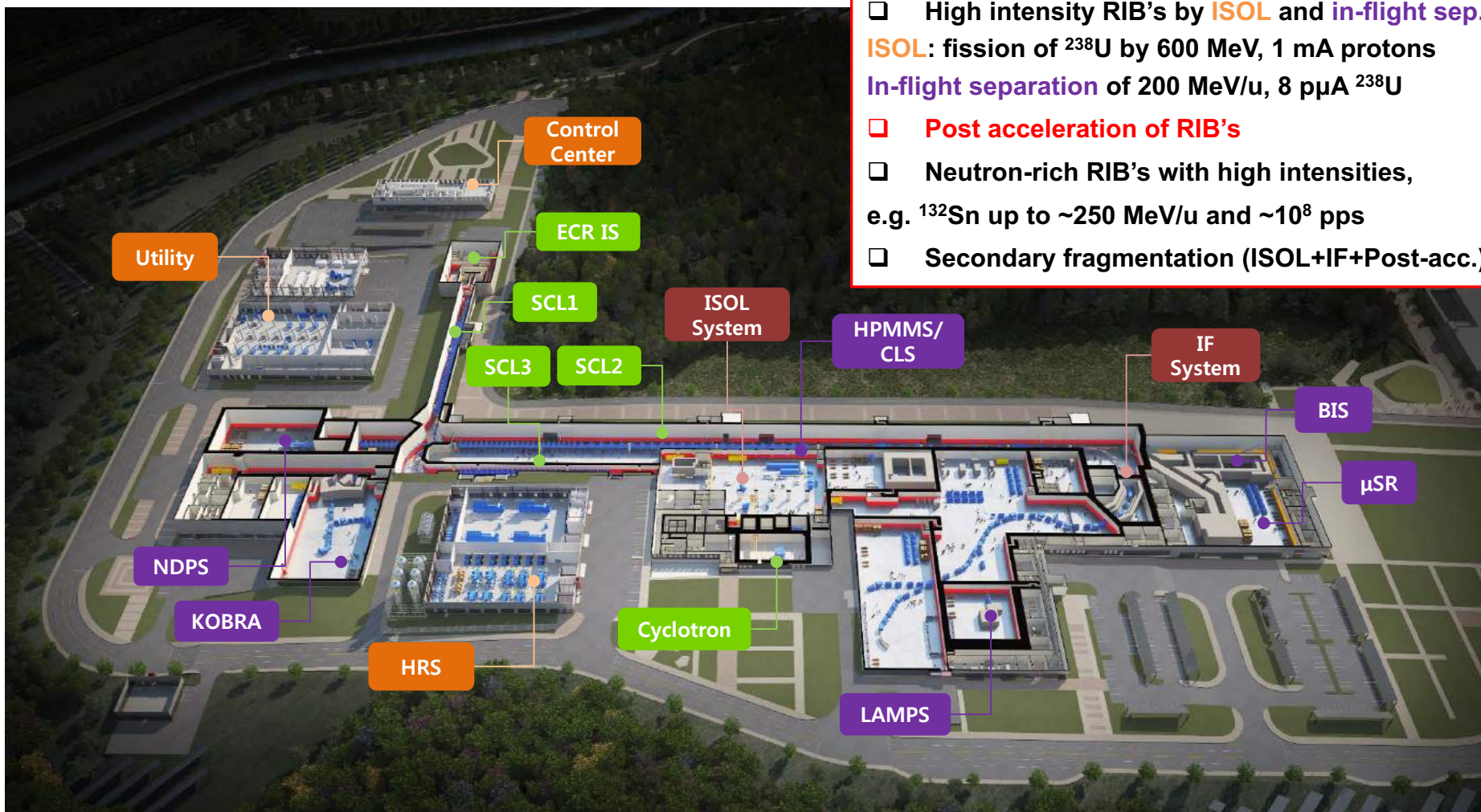


> 16 existing RIB facilities
 > 7 facilities under construction
 > 8 Projects – design & R&D

3 injectors + cascade of 4 cyclotrons
 ⇒ several to 345 MeV/nucleon
 A variety of primary beams
 World highest-intensity RI beams



≥ 500 M\$, operation started in 2007
 40M\$ upgrade is ongoing -> intensity x30



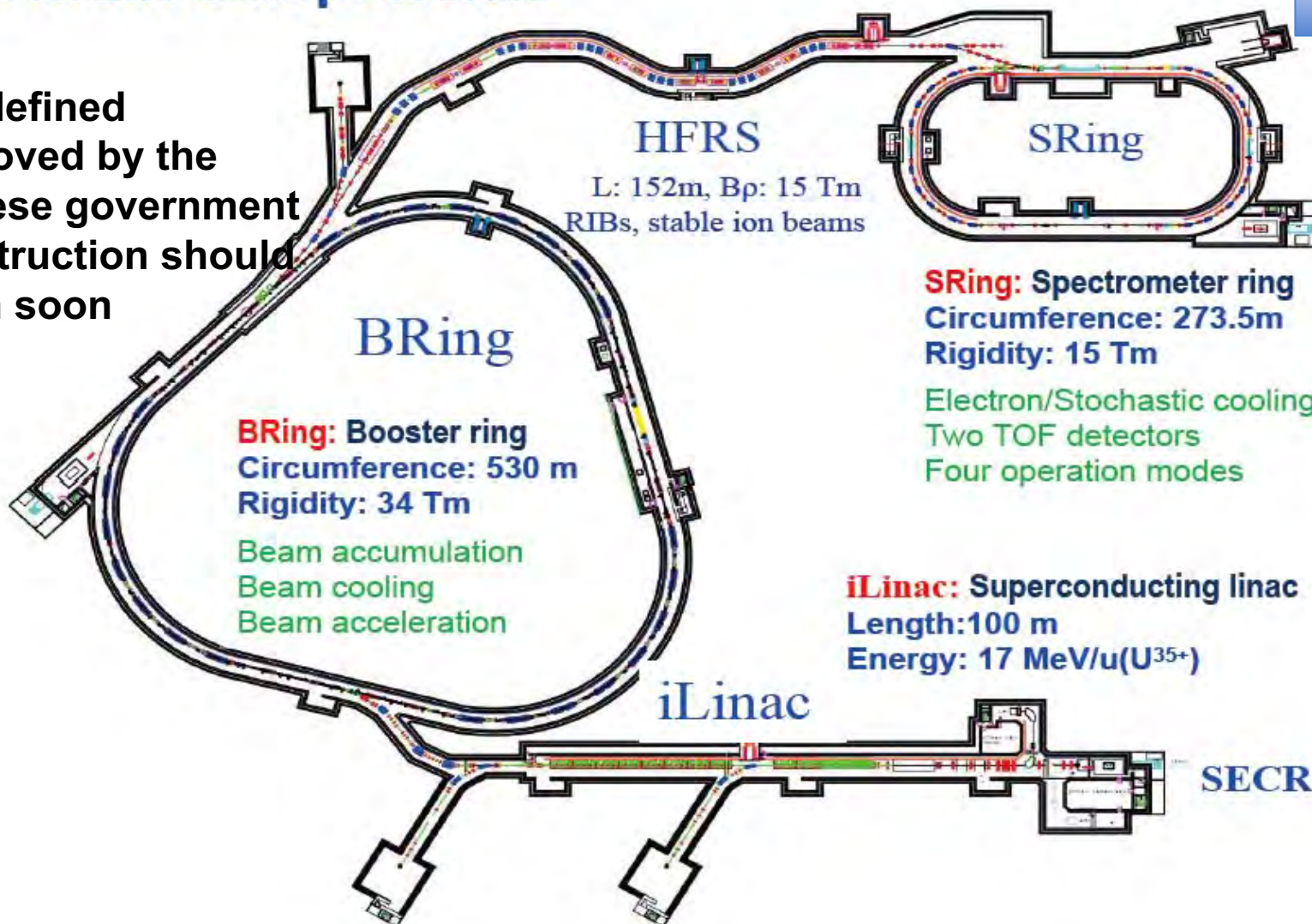
- ❑ High intensity RIB's by **ISOL** and **in-flight sep.**
ISOL: fission of ^{238}U by 600 MeV, 1 mA protons
In-flight separation of 200 MeV/u, 8 μA ^{238}U
- ❑ **Post acceleration of RIB's**
- ❑ Neutron-rich RIB's with high intensities, e.g. ^{132}Sn up to ~ 250 MeV/u and $\sim 10^8$ pps
- ❑ Secondary fragmentation (ISOL+IF+Post-acc.)

≥ 1 B\$, construction started in 2017 and to be finished by 2023

The main components

in China

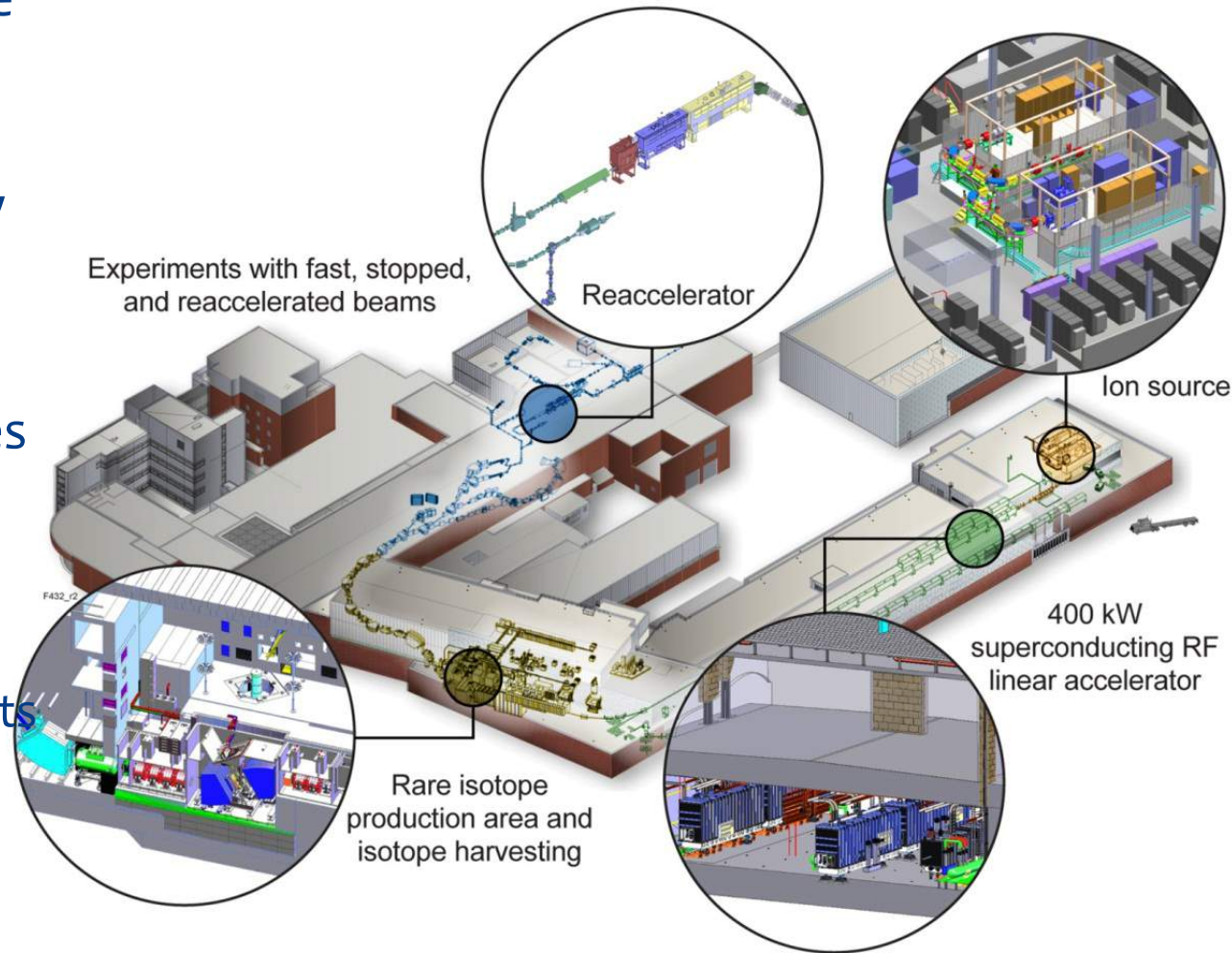
- Site defined
- Approved by the Chinese government
- Construction should begin soon

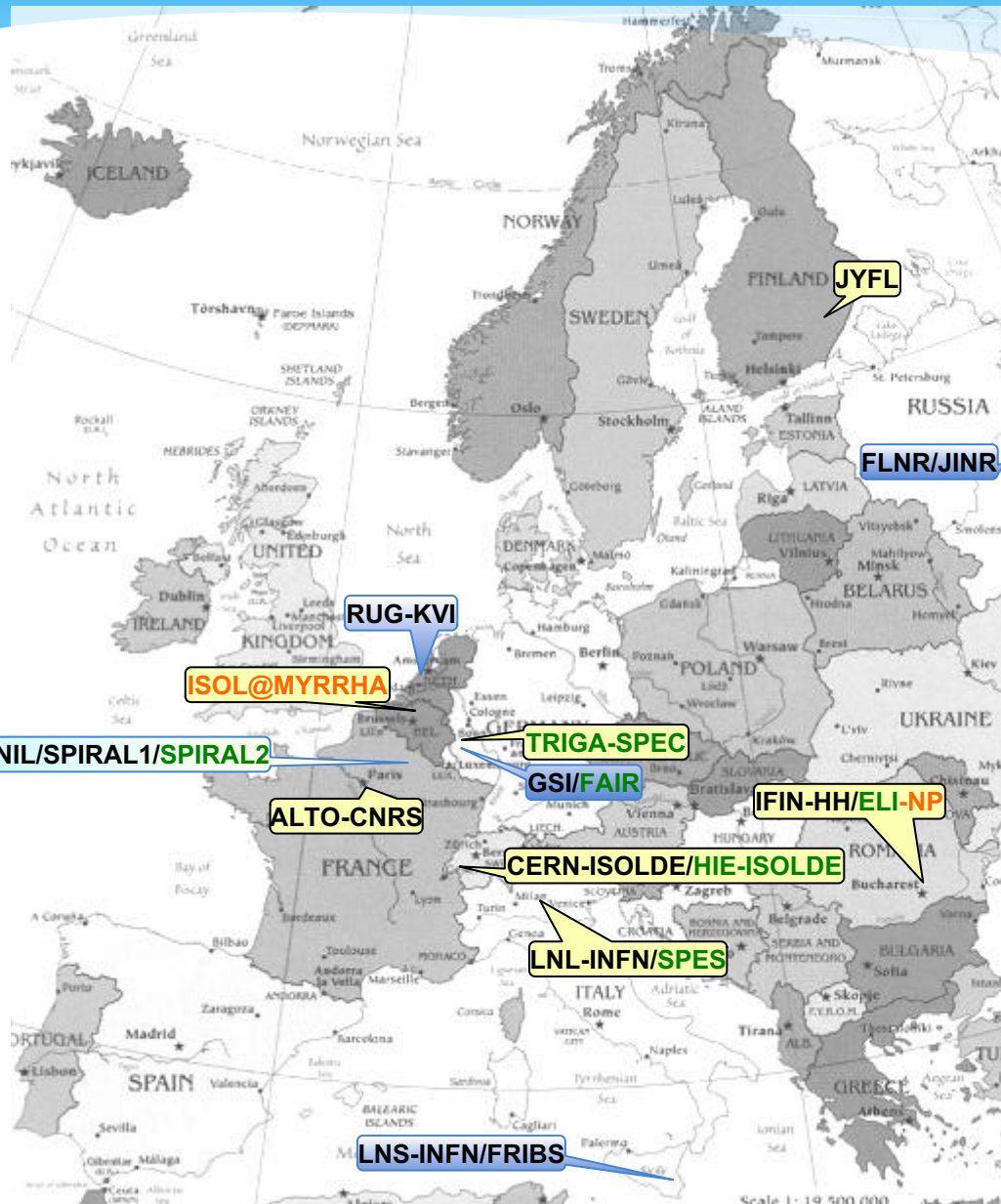


≥ 1 B\$, site identified, construction will start in 2018

Major US Project – Facility for Rare Isotope Beams, FRIB

- * Funded by DOE Office of Science – 2020 completion
- * Key Feature is 400kW beam power (200 MeV/u 5×10^{13} $^{238}\text{U}/\text{s}$)
- * Separation of isotopes in-flight
 - * Fast development time for any isotope
 - * Suited for all elements and short half-lives
- * Reaccelerated beams up to 12 MeV/u





9 Existing RIB Facilities:

5 In-flight fragmentation

4 ISOL

5 Facilities/upgrades under construction or commissioning

2 Projects under design

Community: 2700-3000 scientists and highly qualified engineers

Transnational access within EU projects



Nuclear structure reactions and applications

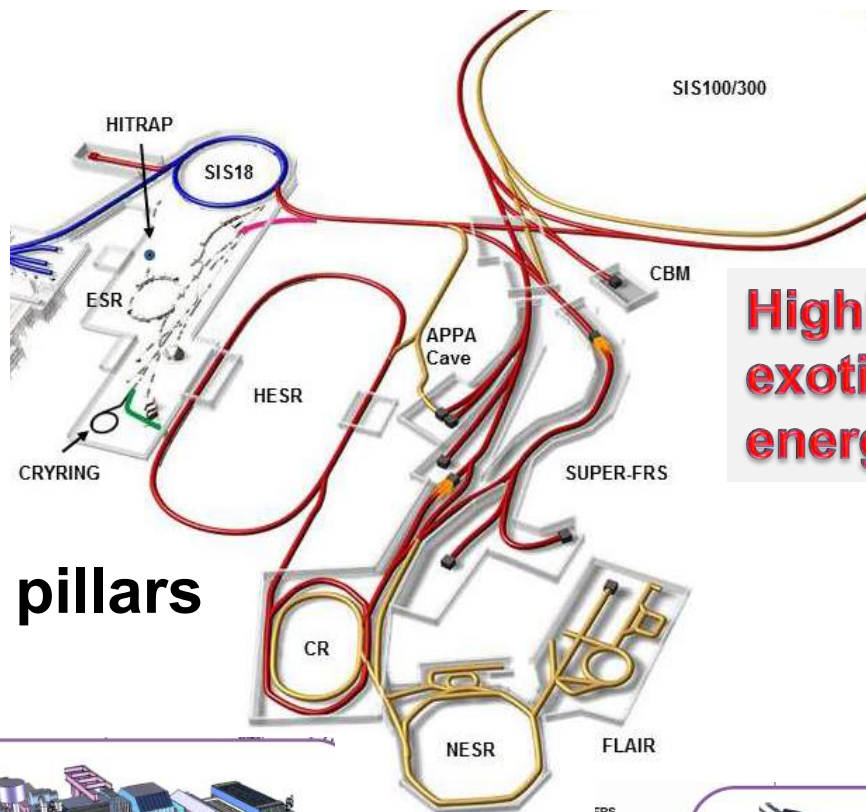
- GANIL (France)
- LNL-LNS (Italy)
- ISOLDE (CERN)
- JYFL (Finland)
- ALTO (CNRS, France)
- GSI (Germany)
- KVI (The Netherlands)
- NLC (HIL/IFJ PAN, Poland)
- IFIN-HH/ELI-NP (Romania)
- ECT* (Italy)



Hadron physics with hadronic and electromagnetic probes

- CERN (LHC, COMPASS, fixed target)
- GSI/FAIR (Germany)
- LNF, Frascati Italy
- MAMI, Mainz Germany
- ECT*, Trento Italy
- ELSA, Bonn Germany
- COSY, Julich Germany

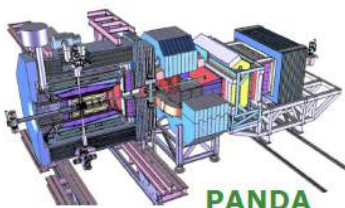




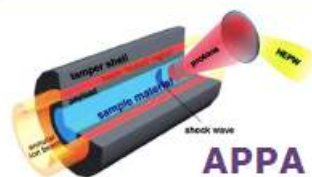
Large facility covering all nuclear physics domains !

Highly charged ions (e.g. U^{92+}) and exotic Nuclei) from rest to relativistic energies 4.9 GeV/A

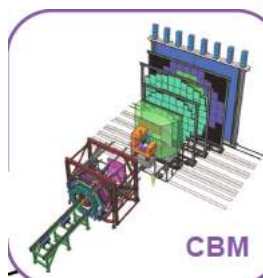
4 pillars



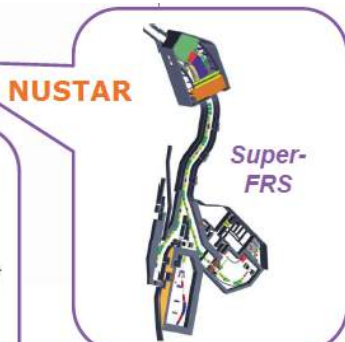
PANDA



APPA

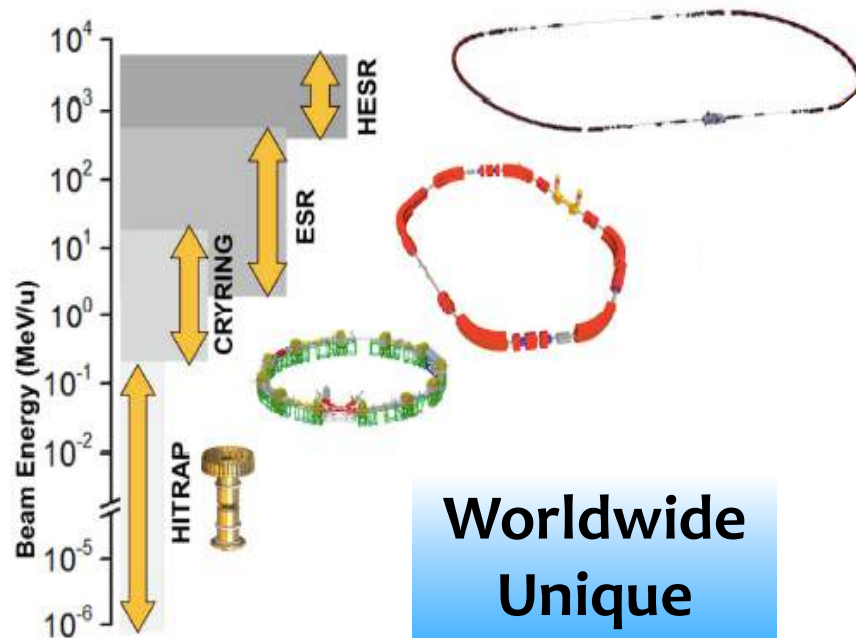


CBM



NUSTAR

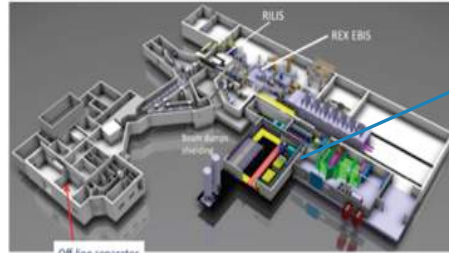
Super-FRS



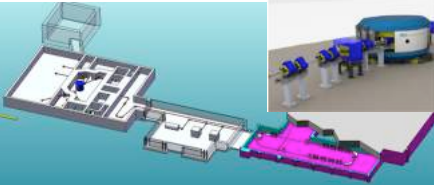
Worldwide Unique

An effort to perform a coordinated program with existing and planned radioactive beams produced with the ISOL technique and to do R&D for the next step EURISOL

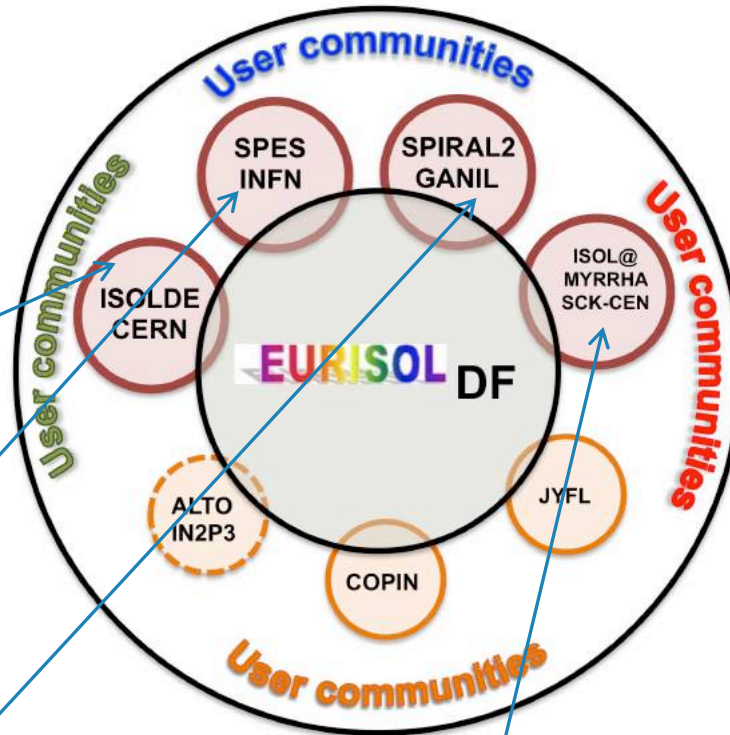
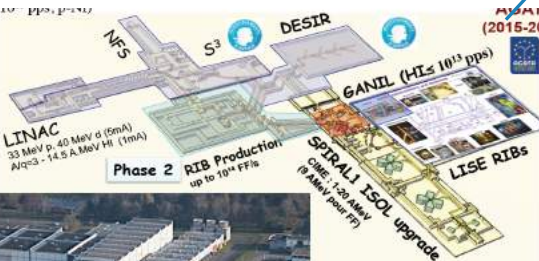
ISOLDE-CERN



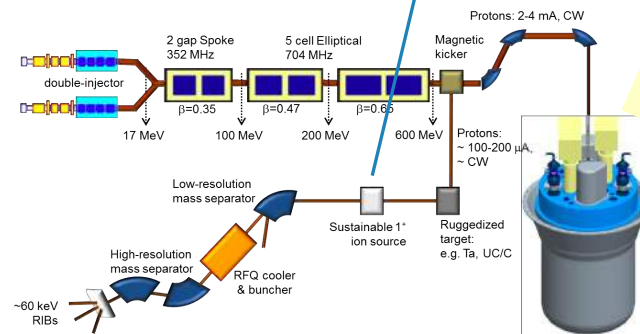
SPES at LNL



GANIL/SPIRAL2-France

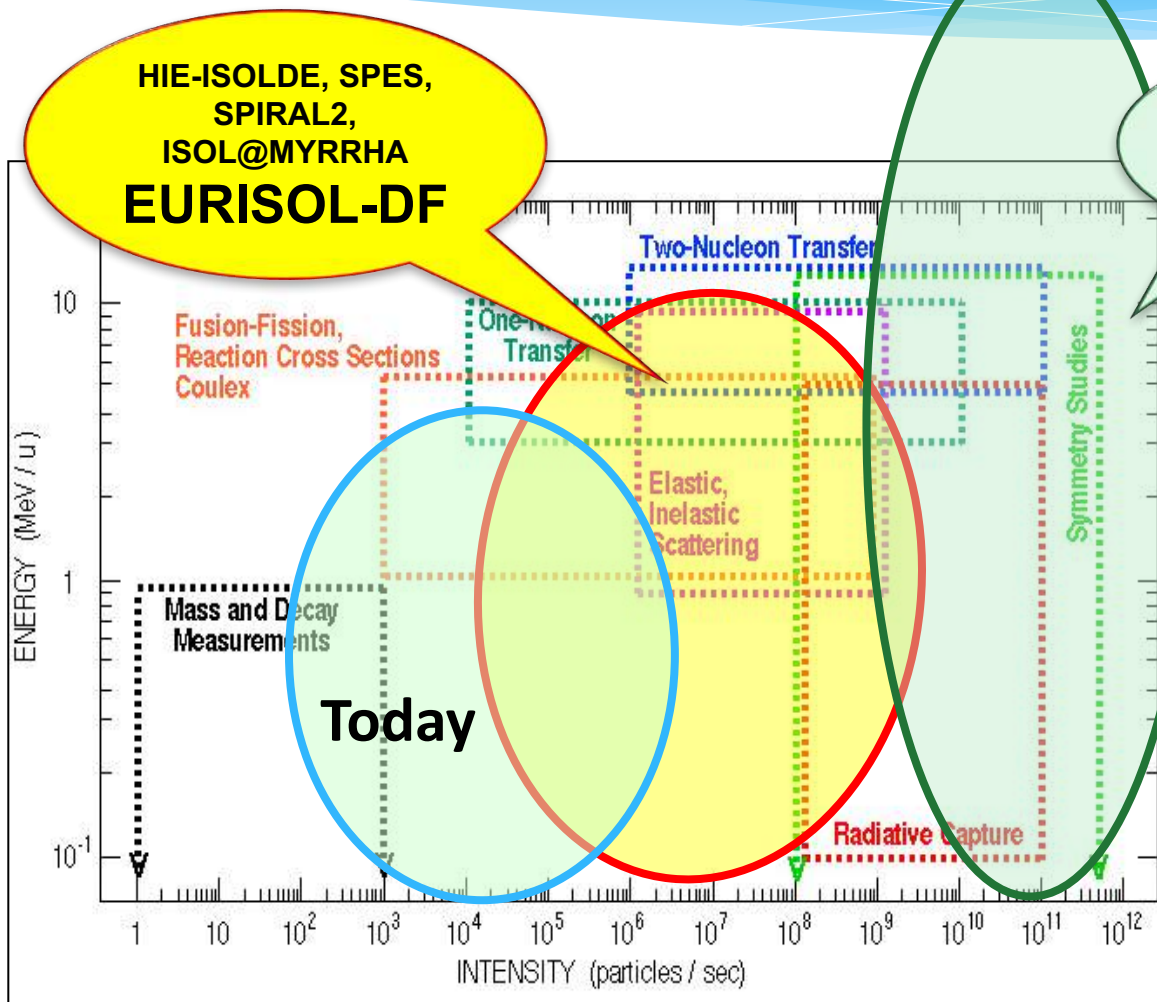
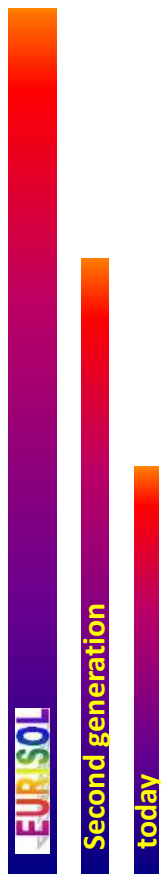


Focus is on nuclear structure, nuclear astrophysics and applications

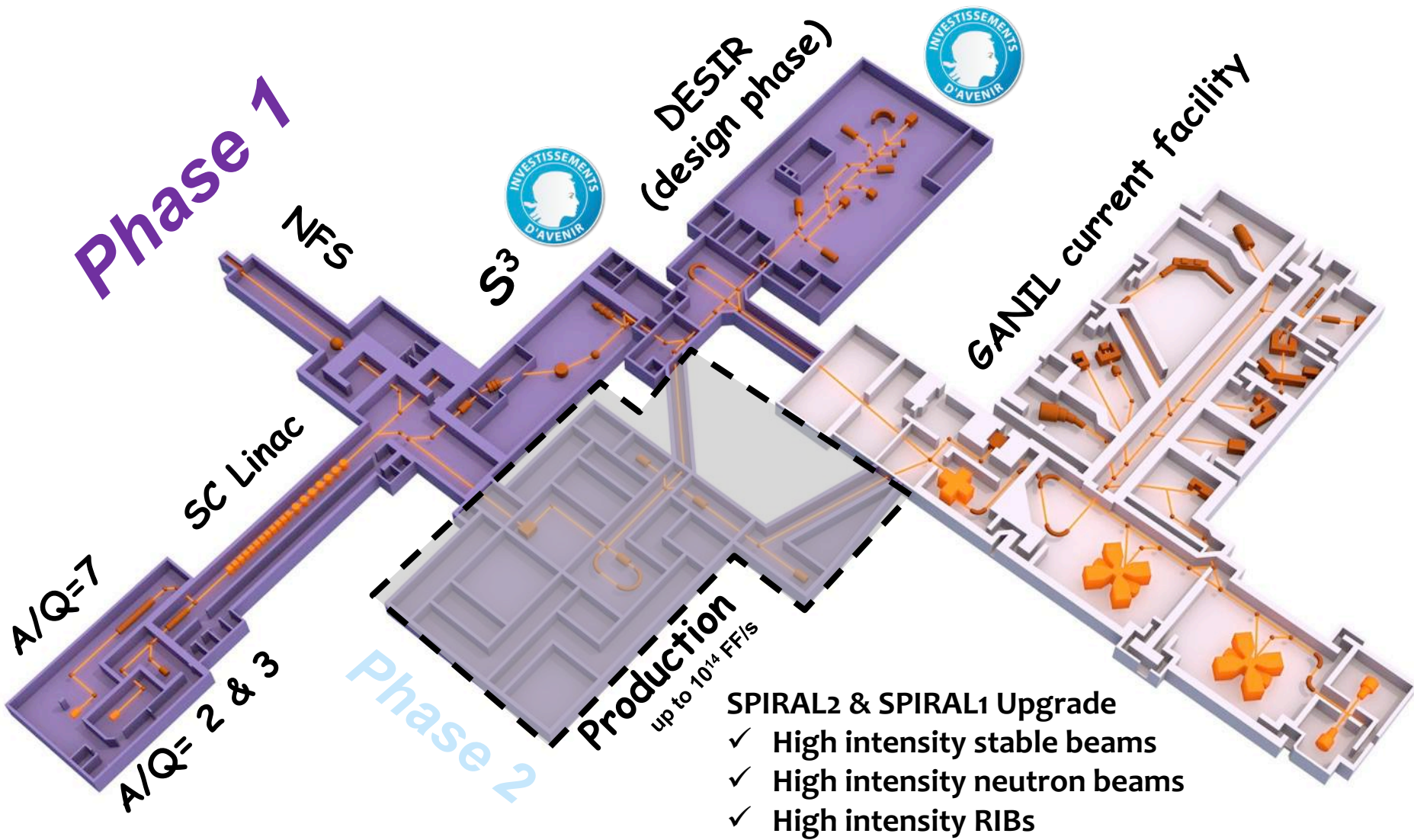


ISOL@MYRRHA-
Radioactive beams
for science and appl.

RIB Energies from 0 to 20MeV/nucl.

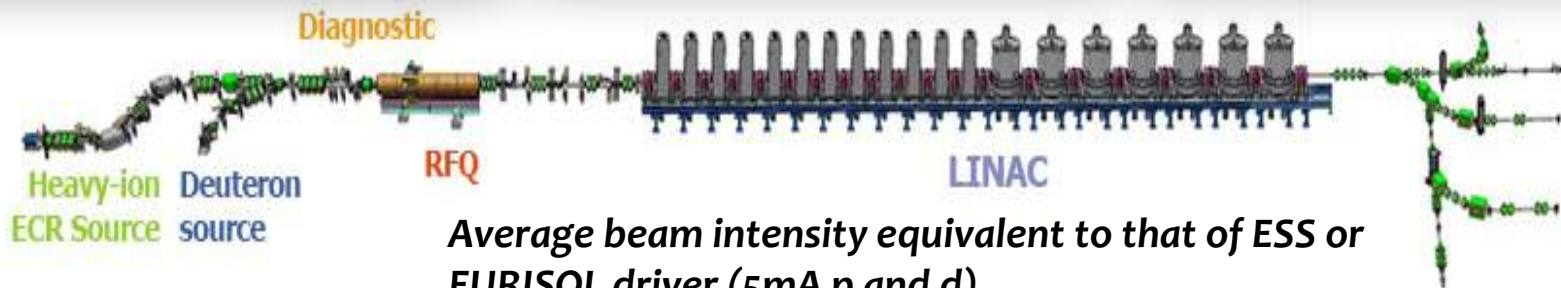
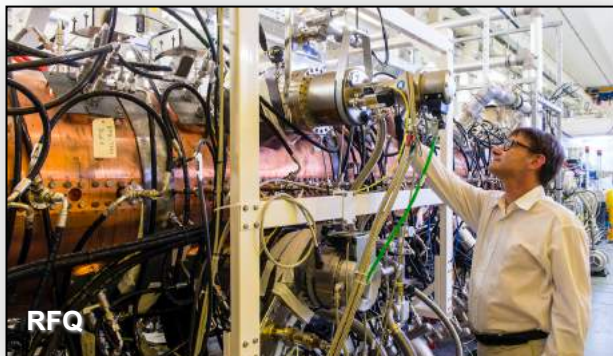


-> EURISOL-DF (Distributed Facility) Initiative from 2014 as an intermediate step towards EURISOL



SPIRAL2 & SPIRAL1 Upgrade

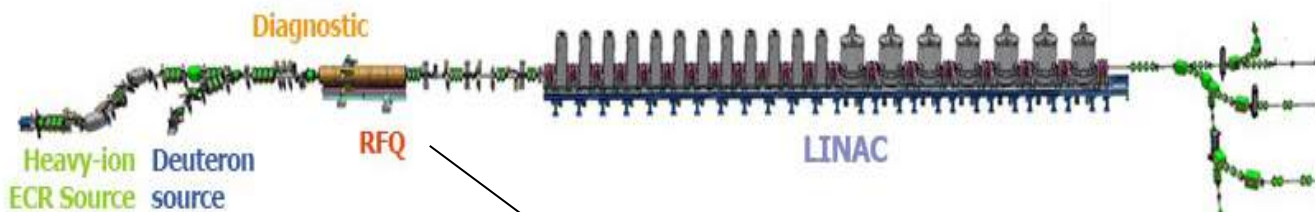
- ✓ High intensity stable beams
- ✓ High intensity neutron beams
- ✓ High intensity RIBs



Average beam intensity equivalent to that of ESS or EURISOL driver (5mA p and d)



Installation is almost complete



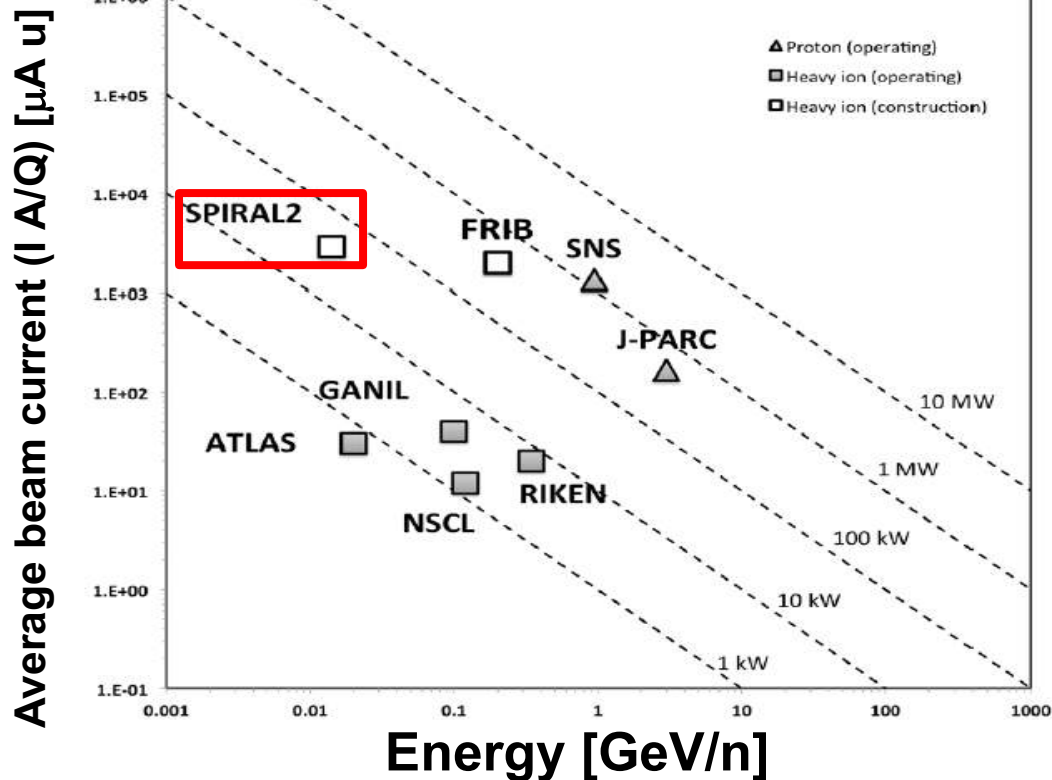
1000 μA $^4\text{He}^{2+}$
 150 μA $^{18}\text{O}^{6+}$
 3,2 μA $^{40}\text{Ar}^{14+}$ (60 kV)

5 mA protons ($Q/A=1$)
 500 μA $^4\text{He}^{2+}$ ($Q/A=1/2$)
 100 μA $^{18}\text{O}^{6+}$ ($Q/A=1/3$)

Ions	Intensity (μA) [A/Q=3]	High Intensity (μA) [A/Q=6-7]
^{18}O	216	375
^{19}F	28,6	50
^{36}Ar	17.5	40
^{40}Ar	2.9	30
^{36}S	4.6	30
^{40}Ca	3	20
^{48}Ca	1.25	15
^{58}Ni	1.1	10
^{84}Kr	0	20
^{124}Sn	0	10
^{139}Xe	0	10
^{238}U	0	2.5

$\times 5-10$

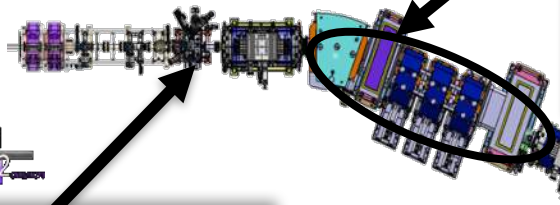
$\times 10^x$



S³

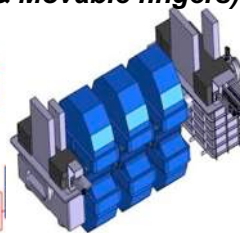
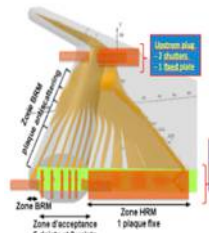
(L=26m)

Beam spot :
 $\sigma_x=0.5\text{mm}$, $\sigma_y[0.5-2.5\text{mm}]$
 Energy precision $\approx 5 \cdot 10^{-3}$

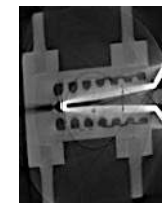


Dispersive zone

(beam dump & Movable fingers)



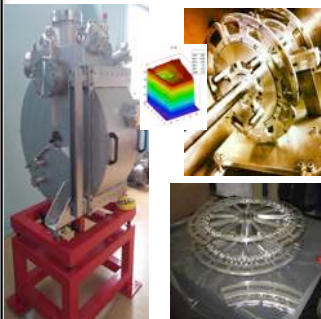
lrfu
 Institut de recherche sur les lois fondamentales de l'Univers



tested for 5kW/cm²

GANIL
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Target system



High power rotating targets (3000-5000 rpm)
 Stable & Actinide systems

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3 x M-dipoles



Large H & V gaps

E-Dipole



20 cm gap & +/- 350 kV
 E_{pmax} : 12-14 MeV
 Open slit in the anode

IPN
 Institut de Recherche Nucléaire
 ORSAY

SC Multipoles



Q+S+O fields



PSS



Cold Box

F. Dechery et al., Eur. Phys. J. A (2015) 51: 66

F. Dechery et al., in press NIMB

**Full assembly & tests at GANIL
 planned in 2019**



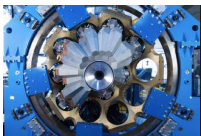
Complete urgently the construction of the ESFRI flagship FAIR and develop and bring into operation the experimental programme of its four scientific pillars APPA, CBM, NUSTAR and PANDA.

Support for construction, augmentation and exploitation of world leading ISOL facilities in Europe towards EURISOL.



Support for the full exploitation of existing and emerging facilities.

Support for ALICE and the heavy-ion programme at the LHC with the planned experimental upgrades.



Support to the completion of AGATA in full geometry.



eurorib 2018

Giens, France

May 27th - June 1st

Topics

*Future RIB facilities
Nuclear astrophysics
Transuranium nuclei
Fundamental interactions
Applications to other fields
Direct Reactions with radioactive beams
Nuclear structure far from stability and hypernuclei
Dynamics and Thermodynamics of exotic nuclear systems
At and beyond the dripline and new modes of radioactivity
Instrumentation, electronics and data acquisition systems
Production and manipulation of RIB*

EURISOL

FAIR

GANIL

GSI

ISOLDE

SPES
exotic beams for science

IDEA
AL

IDEA AL
ERDF is funding the
European Union research and
innovation programme
under grant agreement No 101019718

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Thank you for
your attention

Dziękuję za uwagę